

# **Effects of Polarisation Fading upon the Performance of Skywave Radars.**

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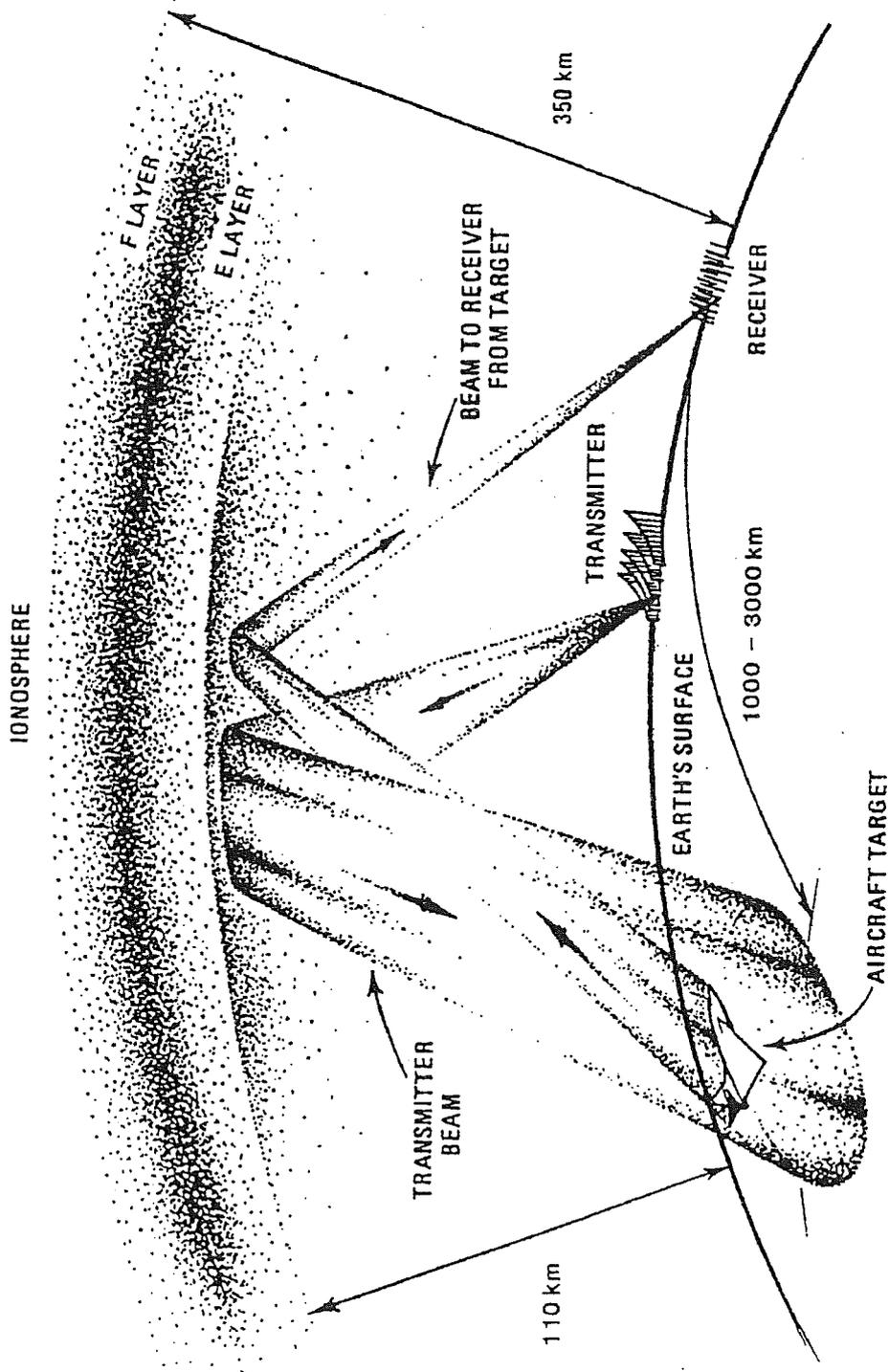


Figure 1.1 Overview of radar geometry.

## 1. INTRODUCTION

During my vacation employment at DSTO Salisbury, I was assigned to examine how Polarisation (Faraday) rotation affects Over the Horizon Radars (OTHRs) and what techniques can be used to reduce them. Like any system which relies upon HF radio waves that travel through the ionosphere, OTHR systems like Australia's Jindalee are susceptible to ionospheric fading, and thus a loss in sample resolution. While the effects of polarisation rotation are examined closely in this paper, other cause of signal fading are outlined with respect to their effects upon the operation of OTHR.

As a target, or in this case a beacon is required before any backscatter fading can be measured, I initially investigated how critical the beacon's design parameters had to be in order to achieve the polarisation invariance required to study polarisation fading. Different designs and signal modulations were also examined to the determine what configuration would provide the greatest detectable return power.

Static models were then used to analyse how polarisation fading behaved as the operational frequencies and target ground ranges were varied and the accuracy of certain approximations were investigated. These results were then compared with ray traced values obtained after travelling ionospheric disturbances were introduced. Geographically, it is assumed in all calculations that the transmit and receive sites are located at Harts range and Mt. Everard respectively, and that the beacon is located 2000 kilometres due north unless otherwise stated.

Many thanks must go to my supervisor Dr Fred Earl for his guiding wisdom, as well as Mike Whittington for his unrivalled expertise in anything computer related, Dr Chris Coleman and everyone in IE Group for what was both an enjoyable and challenging twelve weeks.

## 2. PASSIVE BEACON

In order to determine the extent to which polarisation fading affects radar performance, and in particular target backscatter across all operational ground ranges, it is proposed to set up a stationary passive beacon which will act as a reflector by reradiating a percentage of the incident radar signal. The two main considerations in the design of the beacon are to maximise retransmission power and achieve polarisation invariance. Note that in this section, all propagation losses are ignored and only the effect of polarisation fading is considered.

### 2.1 SIGNAL MODULATION

Since the beacon is intended for stationary deployment, it lacks the Doppler shift required to uniquely identify its backscatter from the clutter received by the radar. It is thus necessary to modulate the reradiated power in such a way as to create a pseudo-Doppler shift. This can be accomplished by modulating the beacon's antenna loads so that the reradiated signal power varies in time. It has been found by Earl [1] that a square wave modulation allows optimal beacon identification, and it can be shown that a unity mark:space ratio maximises the detectable return power. Given the Fourier series transform of a square wave (equation 2.1) with duty cycle  $\alpha$  and peak amplitude  $V$ , we have:

$$v(\theta) = V \left[ \alpha + \sum_{n=1}^{\infty} \frac{2 \sin n\alpha\pi}{n\pi} \cos n\theta \right] \quad (2.1)$$

Integrating over the square wave's period to get the average AC voltage, we have:

$$\begin{aligned} V_{avg} &= \frac{1}{\pi} \int_{-\frac{\pi}{2}}^{\frac{\pi}{2}} \|v(\theta)\| d\theta \\ &= \frac{V}{\pi} \int_{-\frac{\pi}{2}}^{\frac{\pi}{2}} \left( \sum_{n=1}^{\infty} \left\| \frac{2 \sin n\alpha\pi}{n\pi} \cos n\theta \right\| \right) d\theta \\ &= \frac{V}{\pi} \left[ \sum_{n=1}^{\infty} \left\| \frac{2 \sin n\alpha\pi}{n^2\pi} \sin n\theta \right\| \right]_{-\frac{\pi}{2}}^{\frac{\pi}{2}} \\ &= \frac{4V}{\pi^2} \sum_{n=1}^{\infty} \left\| \frac{4 \sin(4n-3)\alpha\pi}{(4n-3)^2\pi^2} \right\| - \left\| \frac{4 \sin(4n-1)\alpha\pi}{(4n-1)^2\pi^2} \right\| \\ \frac{\partial V_{avg}}{\partial \alpha} &= \frac{4V}{\pi} \sum_{n=1}^{\infty} \left\| \frac{4 \cos(4n-3)\alpha\pi}{4n-3} \right\| - \left\| \frac{4 \cos(4n-1)\alpha\pi}{4n-1} \right\| \\ &\equiv 0 \Leftrightarrow \alpha = \frac{1}{2} \end{aligned} \quad (2.2)$$

One consequence of modulating the retransmitted signal from a stationary beacon is the decrease in average power with respect to a moving target. Given the optimum 50% duty cycle ( $\alpha=0.5$ ) square wave modulation and a voltage separation  $R$  as shown below:

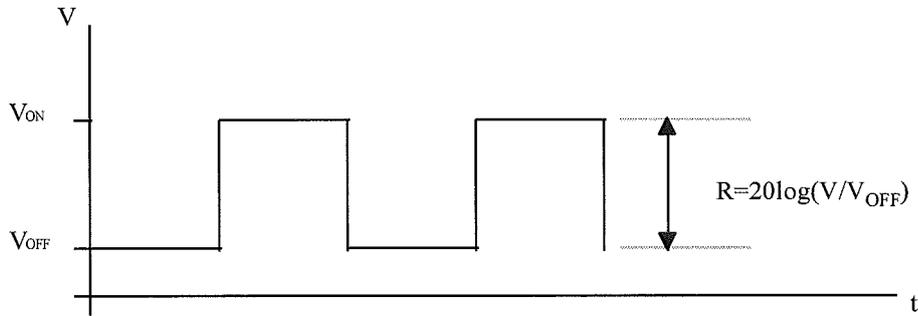


Figure 2.1. Beacon's retransmitted EMF after modulation.

we can see that the amplitude of the first harmonic ( $n=1$  in the Fourier series expansion) is given by:

$$V_{ON} = \frac{V_{max}}{\pi} \qquad V_{OFF} = V_{ON} \times 10^{\frac{-R}{20}}$$

If we define the modulation loss to be the signal loss of the stationary beacon with respect to a moving target, whose signal strength is  $V_{max}$ , we obtain:

$$Modulation Loss = 20 \log \left( \frac{1 - 10^{\frac{-R}{20}}}{\pi} \right) \qquad (2.3)$$

The modulation loss as shown in Graph 2.1 is plotted as  $R$  goes from 0 to 40dB. While the modulation loss tends to 9.91dB as the separation  $R$  approaches infinity, so long as the separation can be maintained at above 20dB, the additional loss will be less than 1dB.

## 2.2 BEACON CONFIGURATION

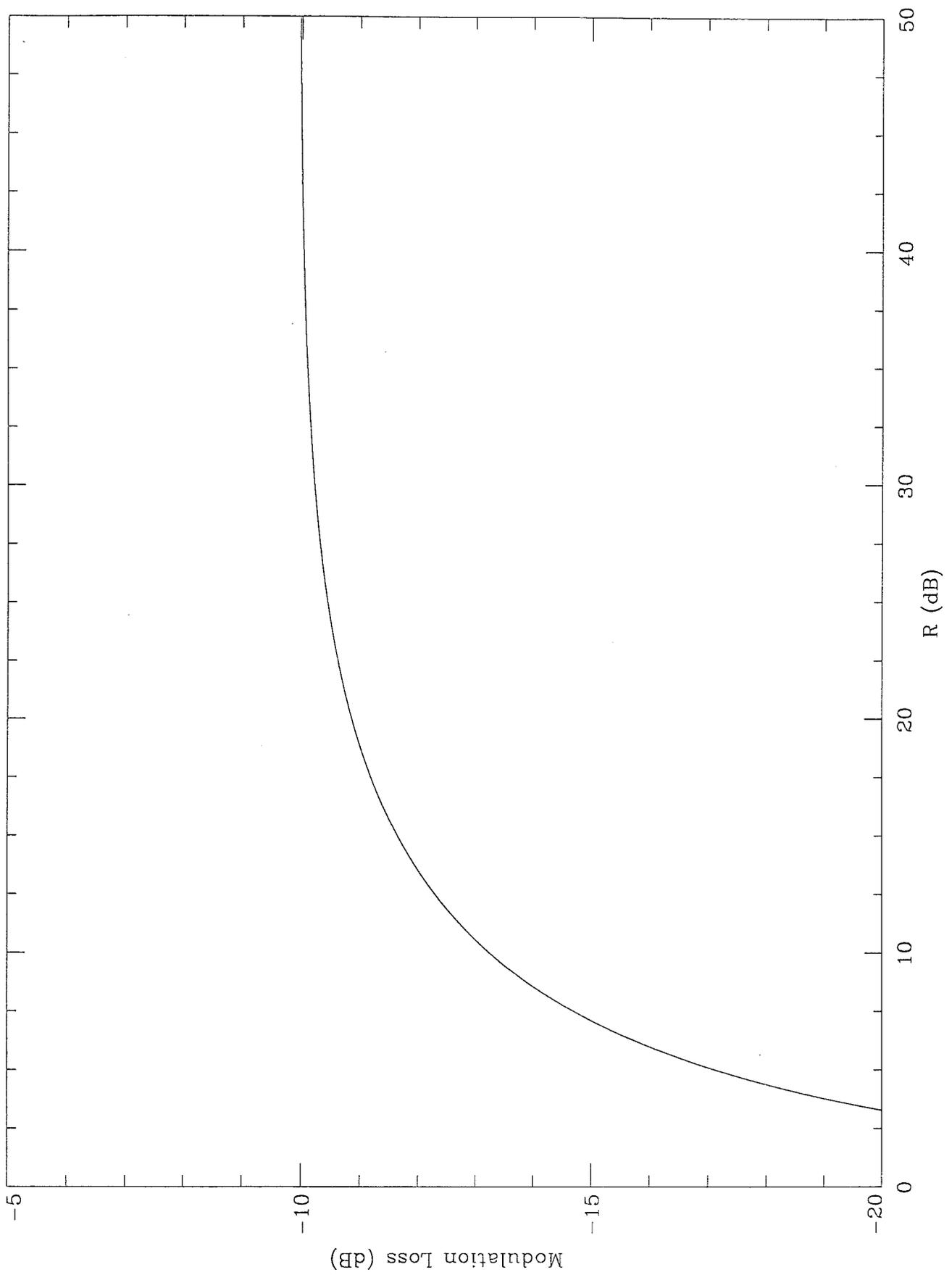
As we are intending to study polarisation, we require the ideal beacon's response to be independent of the angle of the incident ray's polarisation, while maximising the backscatter's signal strength. The three beacon configurations originally considered consisted of:

1. a single vertically polarised antenna
2. a horizontally and a vertically polarised antenna, electrically uncoupled, with a voltage gain ratio

$$\beta_B = \frac{G_H}{G_V}$$

3. a horizontally and a vertically polarised antenna, coupled by a transmission line with phase delay

$$\alpha \text{ and a voltage gain ratio } \beta_B = \frac{G_H}{G_V}$$



Graph 2.1 Modulation losses as a function of voltage separation.

It can be shown that for the above cases, the received EMF as a function of the transmitted EMF  $E$ , polarisation rotation on the outbound leg  $\theta$  and rotation on the inbound leg  $\phi$  is given by:

$$E_R = E \cos\theta \cos\phi \quad (2.4a)$$

$$E_R = E(\cos\theta \cos\phi + \beta_B^2 \sin\theta \sin\phi) \quad (2.4b)$$

$$E_R = \frac{E}{2}(\cos\theta \cos\phi - \beta_B \cos\theta \sin\phi - \beta_B \sin\theta \cos\phi + \beta_B^2 \sin\theta \sin\phi) \quad (2.4c)$$

(see Appendix A for a complete derivation). For the passive beacon to be considered polarisation invariant under planar propagation, at least two orthogonal antennas are required, thereby eliminating case 1 whose retransmitted power would drop to zero as the incident rays approach horizontal polarisation. While both cases 2 and 3 use horizontally and vertically polarised antennas, in case 3 where the two antennas are coupled, we have a situation approximating maximum power transfer as the impedances of the two antennas are assumed to be similar. Thus under ideal circumstances, each antenna is driving a matched load and thus loses half its available voltage (or 75% of its power). As can be seen in Graph 2.4, the coupling of the horizontal and vertical antennas thus greatly affects the reradiated power as a function of polarisation rotation.

To give a qualitative idea of how each beacon configuration performs over all possible polarisation angles, plots of the total received EMF assuming no propagation losses and perfectly matched antennas ( $\beta_B=1$ ) are shown for  $\phi$ ,  $\theta \in [0, 2\pi]$ . As can be seen from Graphs 2.2 to 2.4, case 2 gives the maximum average reradiated power over all polarisation angles, and it is this model that is used hereafter.

For the case with the two uncoupled antennas, the sensitivity of the retransmitted power to variations in the voltage gain ratio  $\beta_B$  was analysed. Normalising the retransmitted power, we find that:

$$P_{Retransmitted} = E^2 [\cos^2 \theta + \beta^4 \sin^2 \theta] \quad (2.5)$$

and to maximise the retransmitted power, we want  $\beta_R$  to approach unity (see Appendix A).

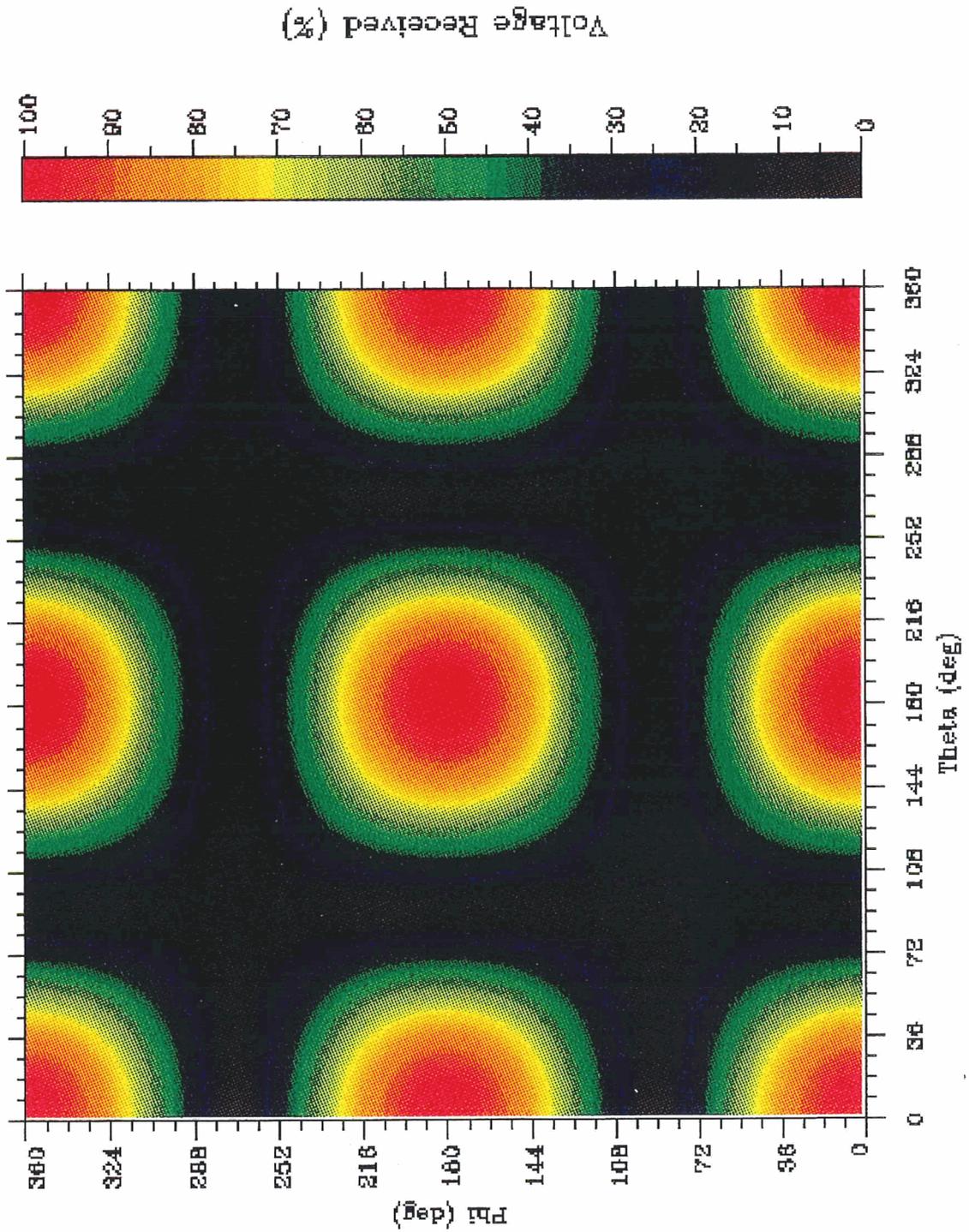
The reradiated power as  $\beta_B$  is varied from 0 to 3dB is shown in Graphs 2.5 through 2.7, and it is clear that by keeping the voltage gains within 1dB of each other, the power loss can be restricted to less than 1.5dB. While it is possible to match two antennas with this degree of accuracy, special care would need to be taken to ensure that the beacon's characteristics remain comparable over the operational frequency range of the radar.

### 2.3 IDENTIFICATION OF BEACON BACKSCATTER

Currently, the receive array at Mt. Everard is only capable of detecting vertically polarised backscatter, thus negating most of the benefits gained from having a polarisation invariant beacon. To allow the polarisation rotation to be studied, it is necessary to measure the horizontal component of any polarised ray.

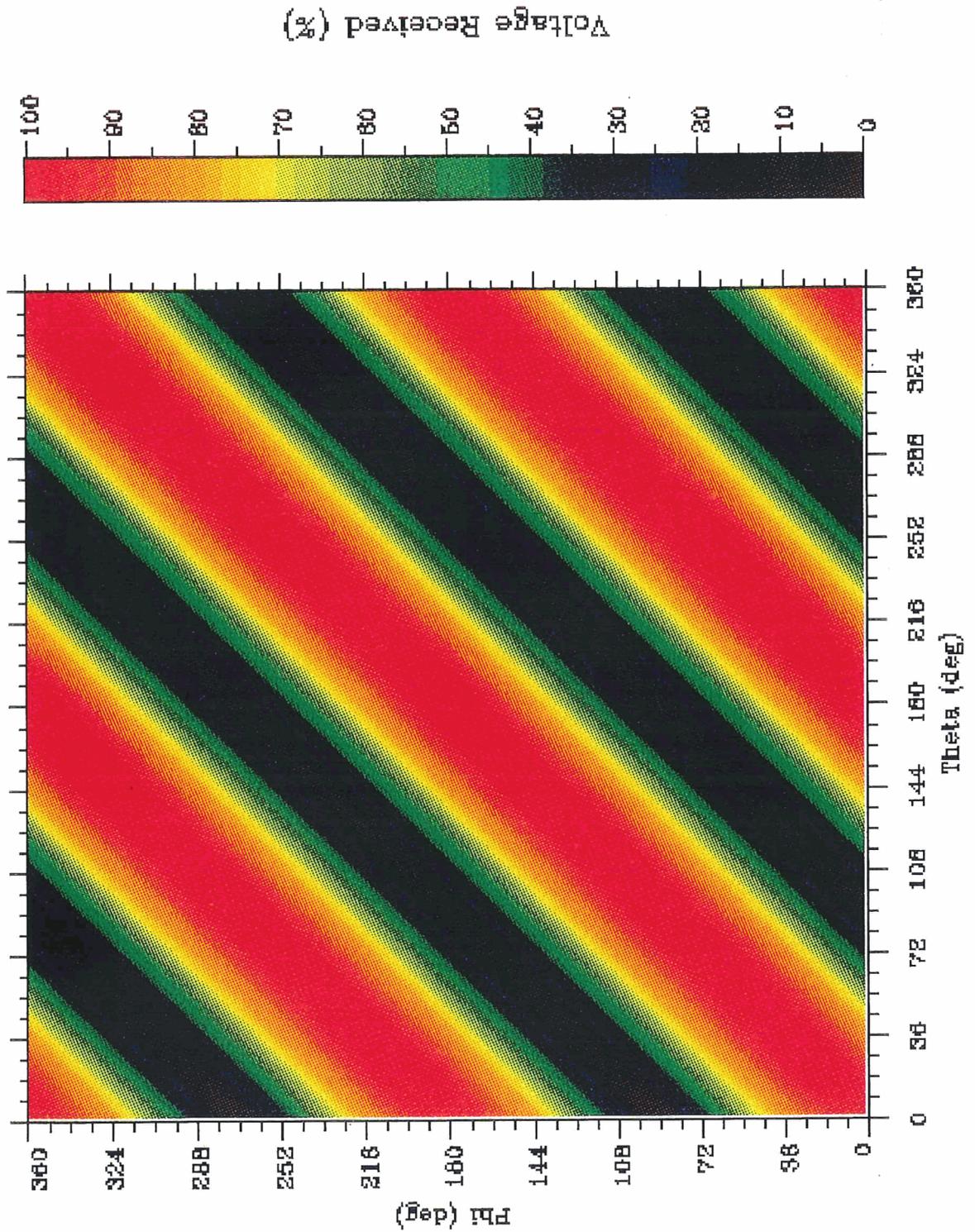
While the previous analysis investigated the dependence of the beacon's polarisation invariance upon  $\beta_B$ , similar reasoning can be applied to determine how closely the horizontally polarised receive antenna needs to be matched to the vertical array before horizontally polarised backscatter can be detected reliably. By extending equation (2.4b) to include this horizontally polarised receive array with a voltage gain ratio  $\beta_R$  defined with respect to the existing vertical array, the expressions for received EMF and normalised power then become:

Ratio of Induced Voltage for Method 1.



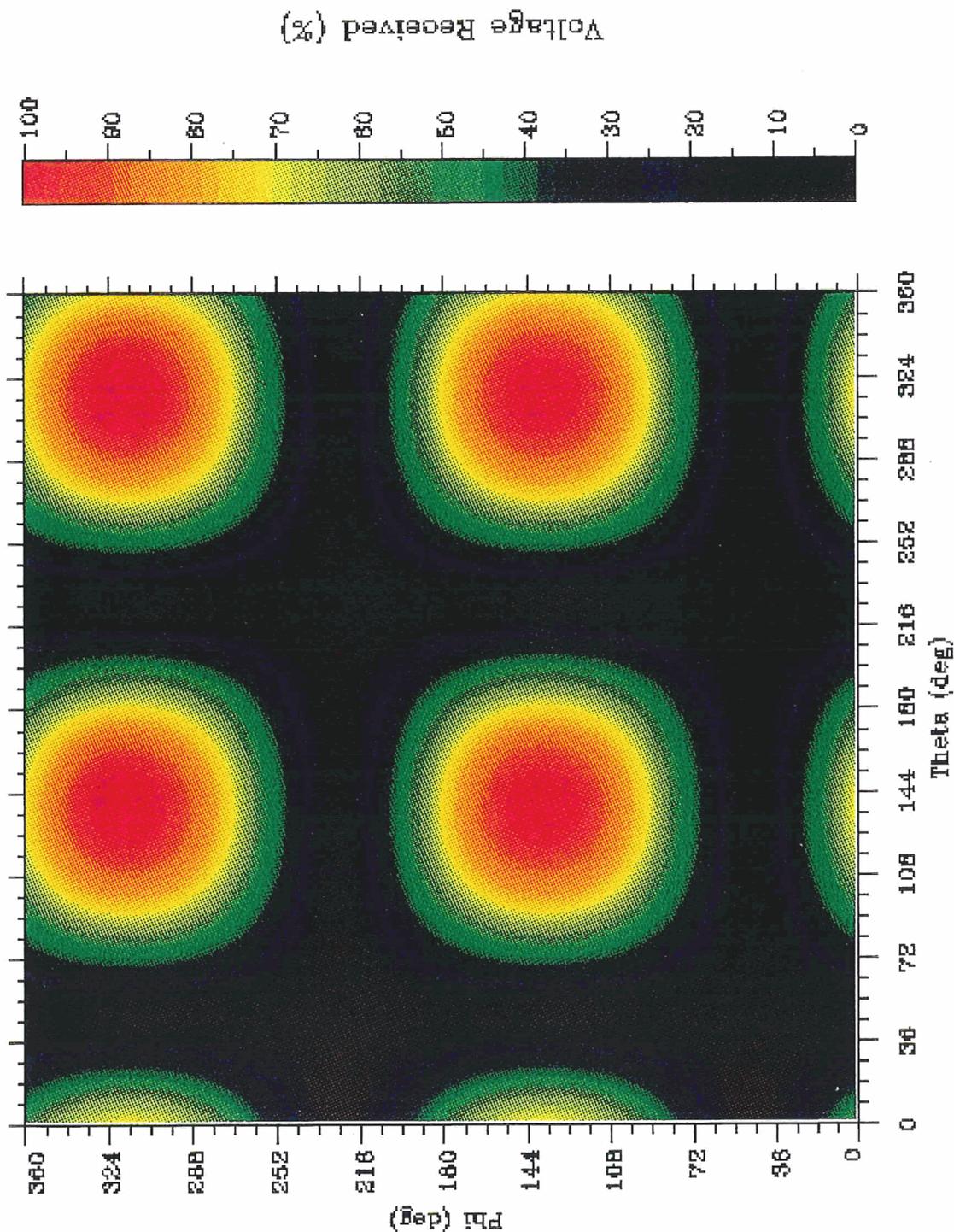
Graph 2.2 Induced voltage ratios for beacon configuration 1.

Ratio of Induced Voltage for Method 2.



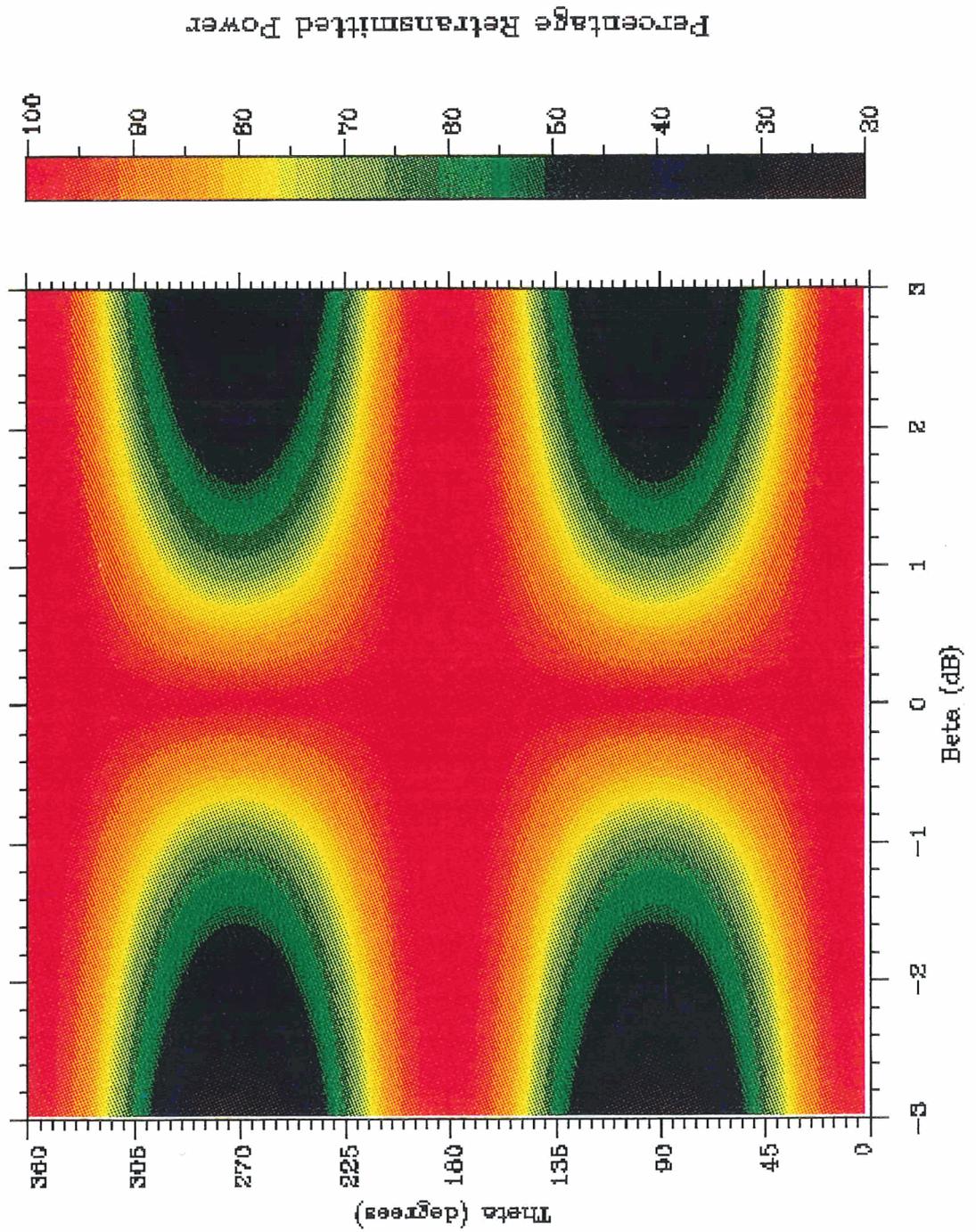
Graph 2.3 Induced voltage ratios for beacon configuration 2.

Ratio of Induced Voltage for Method 3.



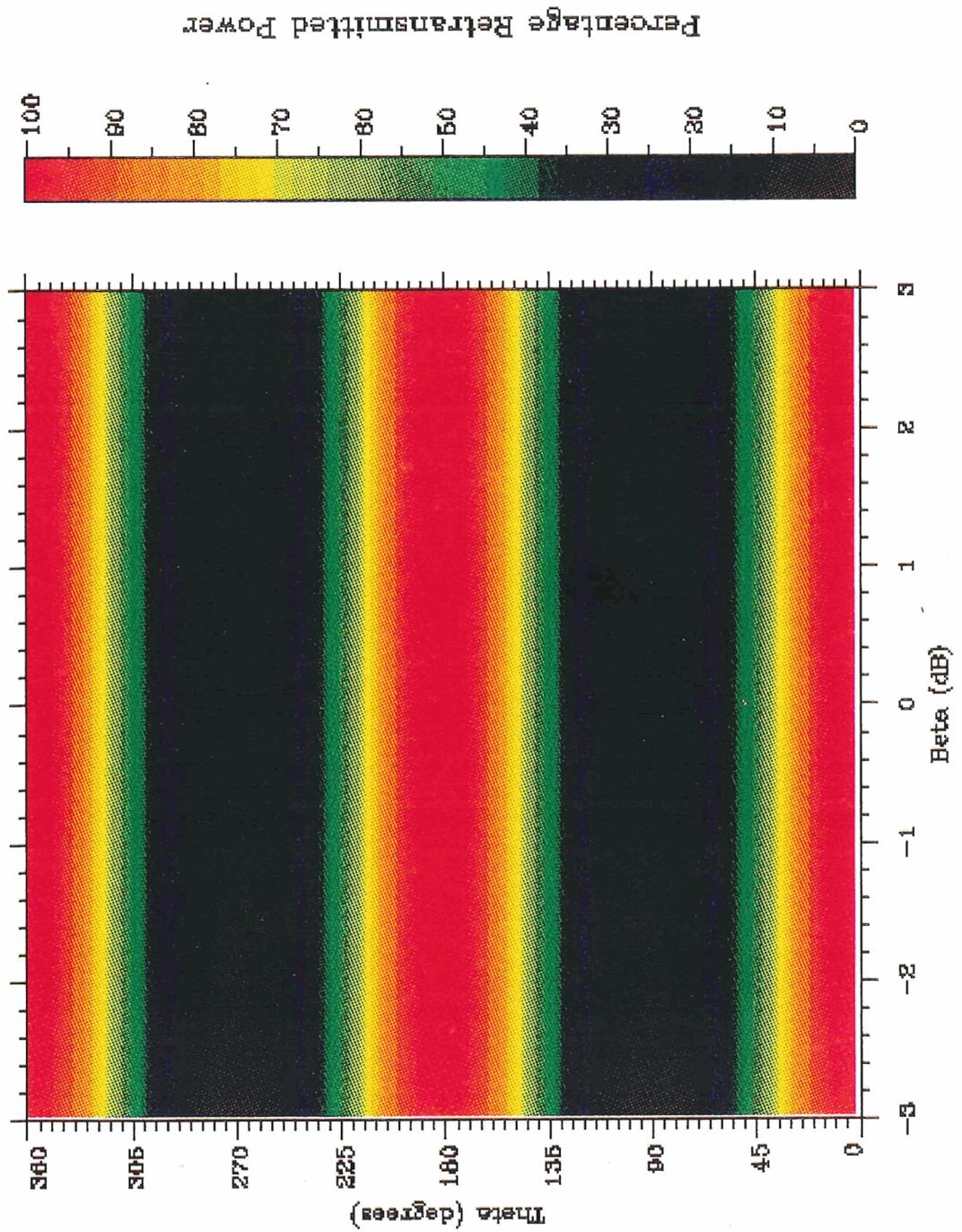
Graph 2.4 Induced voltage ratios for beacon configuration 3.

Total Received Power at Beacon



Graph 2.5 Total received power as a function of  $\beta_R$ .

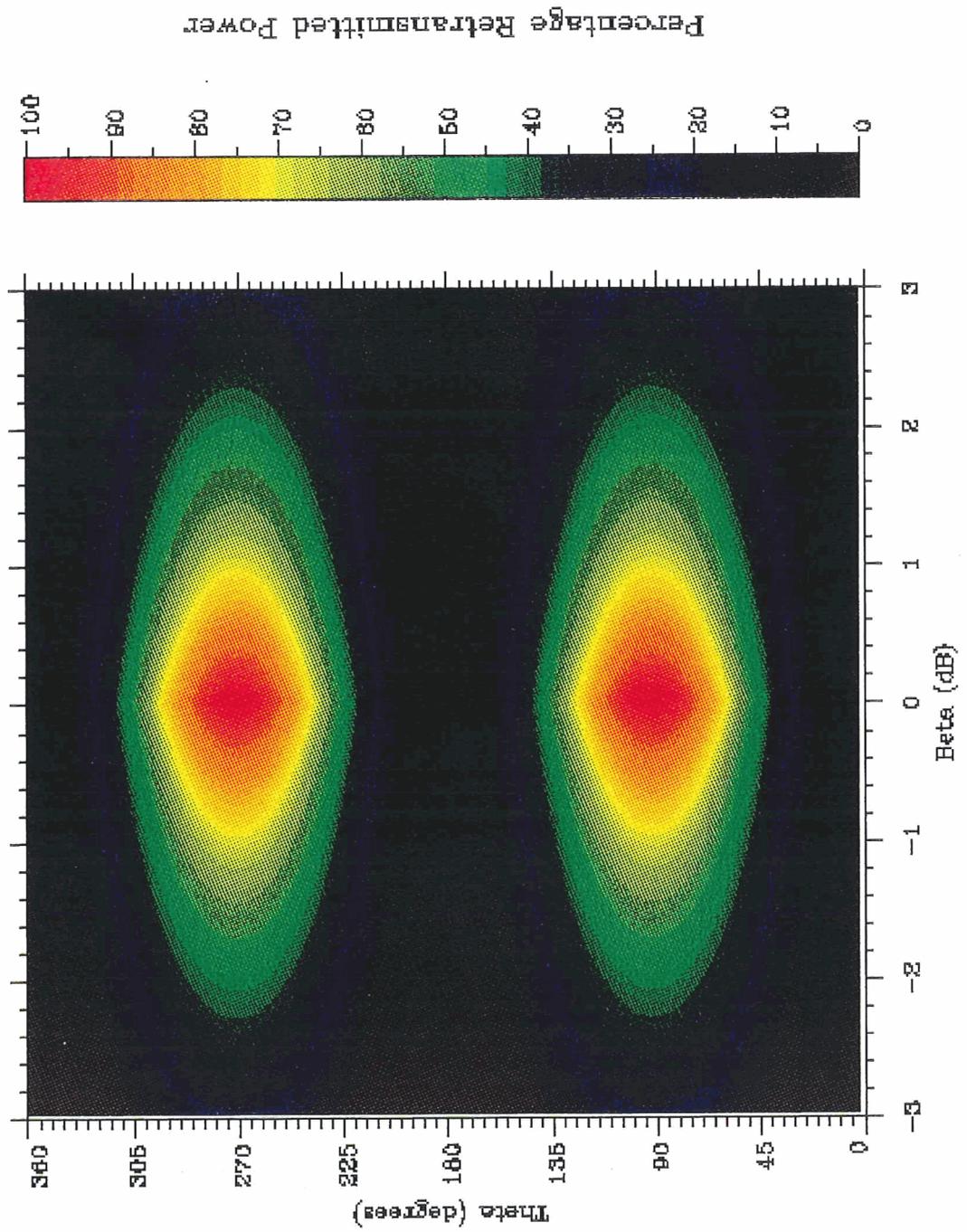
# Power Received by Vertical Beacon Antenna



Graph 2.6

Vertical component of the received power as a function of  $\beta_R$ .

Power Received by Horizontal Beacon Antenna



Graph 2.7

Horizontal component of the received power as a function of  $\beta_R$ .

$$E_{R,Vertical} = E \left[ \cos\theta \cos\phi + \beta_B^2 \sin\theta \sin\phi \right] \quad (2.6a)$$

$$E_{R,Horizontal} = \beta_R E \left[ \beta_B^2 \sin\theta \cos\phi - \cos\theta \sin\phi \right] \quad (2.6b)$$

$$P_R = E^2 \left[ \cos^2 \theta (\cos^2 \phi - \beta_B^2 \sin^2 \phi) + \beta_B^2 \sin^2 \theta (\sin^2 \phi + \beta_B^2 \beta_R^2 \cos^2 \phi) \right] \quad (2.6c)$$

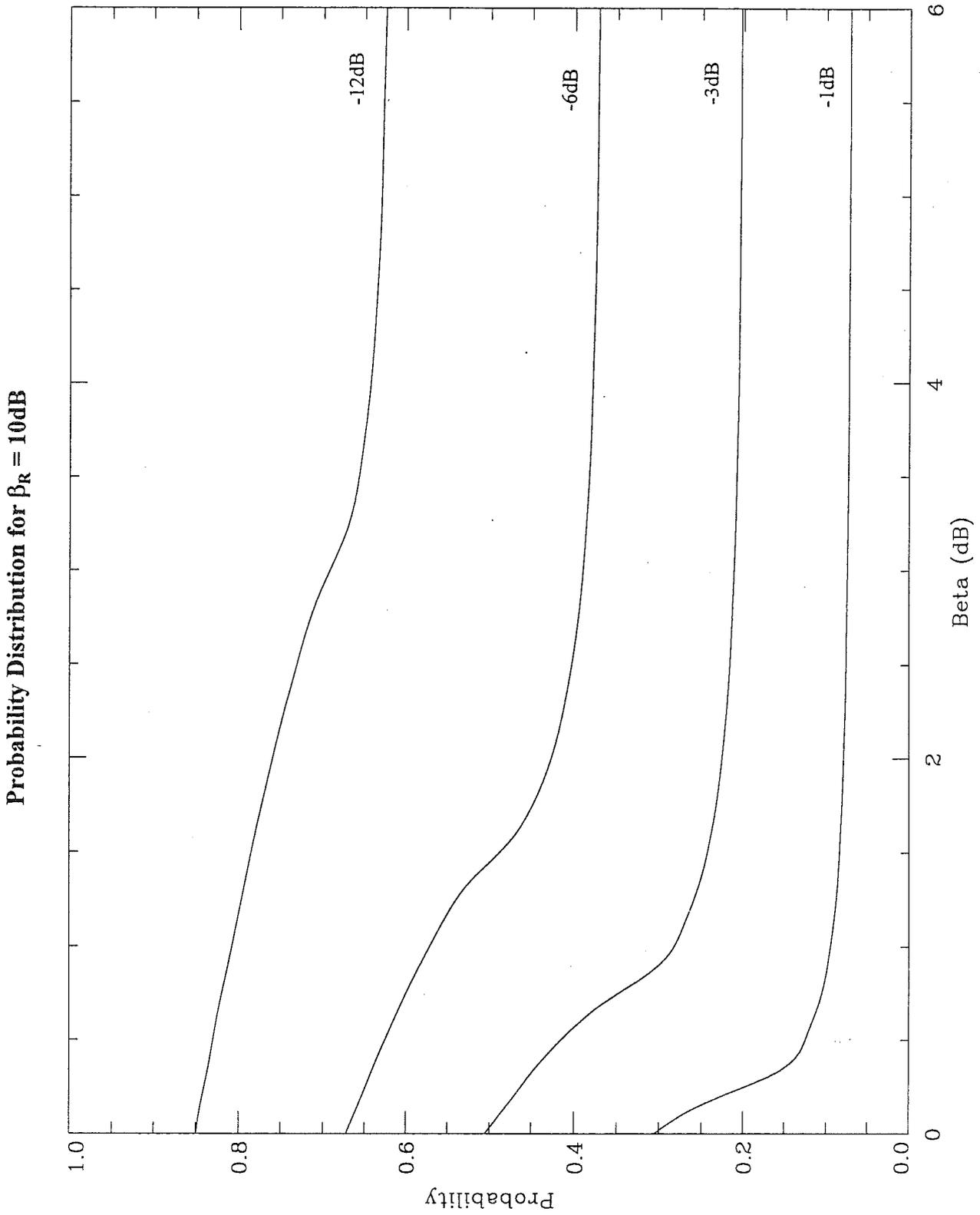
The probability of detecting target backscatter at specific levels below peak signal intensity was calculated by finding the ratio of the number of  $(\phi, \theta)$  values which gave a received signal greater than the desired level to the total pairs of  $(\phi, \theta)$  for specific values of  $\beta_B$ . Probability distributions as  $\beta_B$  varies from 0 to 6dB are shown in Graphs 2.8 and 2.9 where  $\beta_R=10\text{dB}$  and  $\beta_R=1\text{dB}$  respectively.

It was found that the effects of the horizontal receive array were negligible while  $\beta_R$  was greater than around 1.5dB, but as its gain approached the vertical array's the invariant to incoming polarisation became significant. Thus to minimise the effects of polarisation fading at the receive site, it would be necessary to install a horizontally polarised antenna of the order of the existing vertical array.

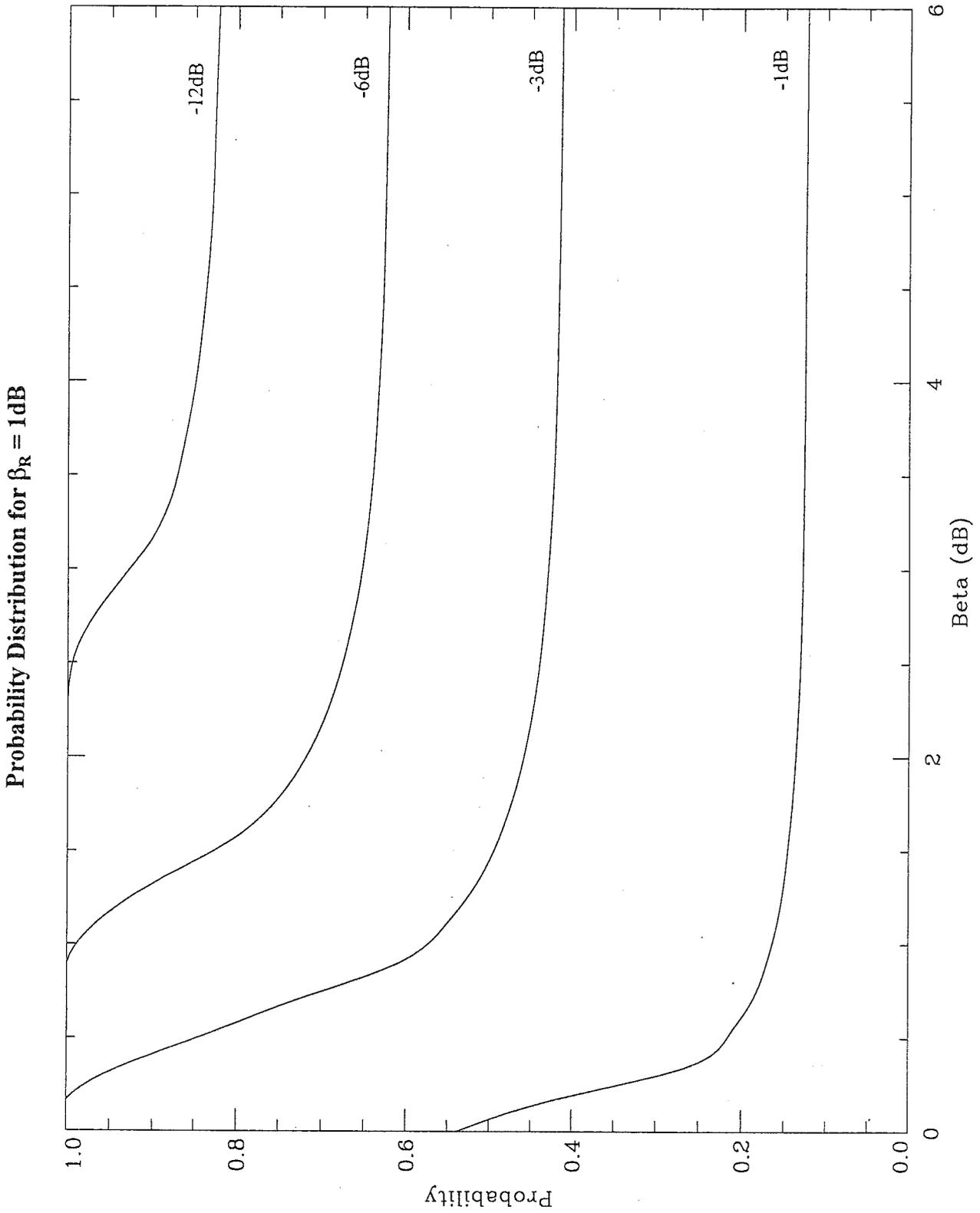
## 2.4 SUMMARY

In closing, it was found that for a system to be considered polarisation invariant, two orthogonal antennas are required, and their voltage gains need to be matched to within 1.5dB. While this is possible for a small structure such as a passive beacon, constructing a horizontally polarised receive array with a gain comparable to the existing vertical array was deemed infeasible, and other techniques to reduce the effects of polarisation fading are examined in the following section.

Also, it was found that while the absolute power loss due to the modulation of the beacon's load tended to 9.91dB as the voltage separation ratio tended to infinity, so long as the separation is maintained at above 20dB, the increase in losses can be kept below 1dB.



Graph 2.8 Signal detection probability distributions for  $\beta_R=10\text{dB}$ .



Graph 2.9 Signal detection probability distributions for  $\beta_R=1\text{dB}$ .

### 3. EFFECTS OF POLARISATION ROTATION

As OTH HF radar systems involve the passage of radio waves through the ionosphere they are vulnerable to the same phenomena that can adversely affect HF communications, including polarisation fading. It should be noted however, that unlike most HF communication systems, polarisation fading is only significant when the incoming signals' ordinary and extraordinary-rays cannot be resolved in group delay. While Faraday rotation provides a good approximation to the polarisation transformation of the ray as it propagates through the ionosphere, there are other causes of polarisation rotation which will not be discussed here.

Polarisation rotation occurs while the ray is in the ionosphere and results in a net change in polarisation angle at the receive site due to the different phase paths of the O and X-rays (for further detail see [3]). Thus under certain circumstances a target may be located in a region whose backscatter, as incident at the receive site, arrives horizontally polarised and is thus undetectable by the radar. As the construction of a large horizontally polarised receive array whose gain is within 1.5dB of the vertical array was determined to be impractical, the use of channel separation and spatial diversity were investigated as methods of minimising the effects of polarisation fading.

While it is desirable to keep the incident ray's polarisation parallel to the receive array, as the received power loss as a function of polarisation angle is given by:

$$P_{dB} = 20 \log |\cos \theta| \quad (3.1)$$

it can be seen from Graph 3.1. that it is not necessary to keep the incident ray vertically polarised, and any system that can restrict  $\theta$  to within  $\pm 60^\circ$  of vertical, will constrain polarisation fading losses to less than 6dB.

#### 3.1 ANALYTIC MODELS

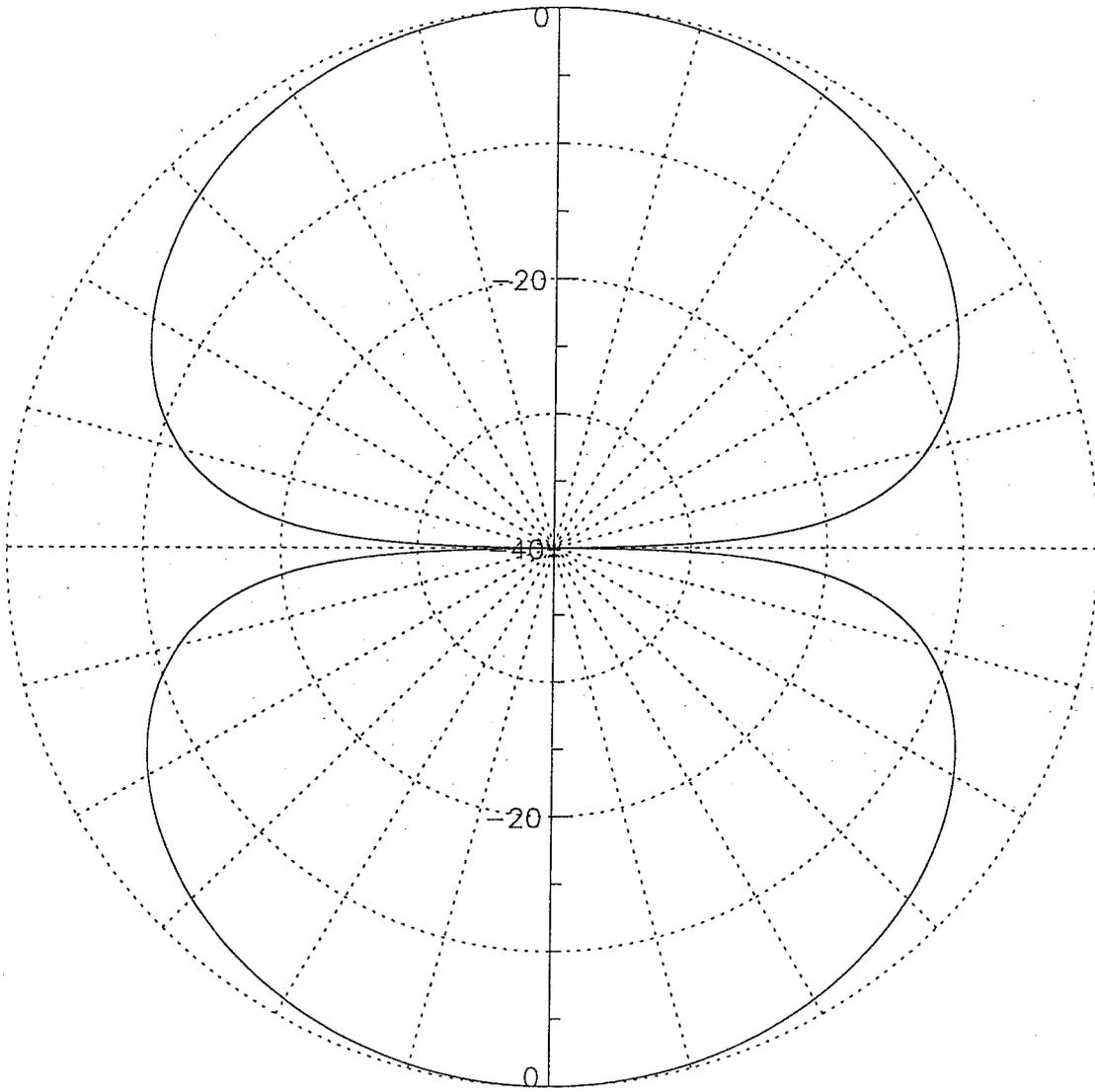
With the exception of the ray traced results, all calculations described hereafter are based upon the exact solutions to the Quasi-Parabolic Ionosphere model described in Croft[2], and assume low-angle propagation for maximum signal strength. As the QP model assumes a homogenous ionosphere, care should be taken to ensure relationships found to be valid in the static model still apply to a dynamic ionospheric model containing TIDs and other ionospheric perturbations.

Values of the ionospheric and physical constants used in all calculations are given in Appendix B, and the equations describing the ray parameters as derived by Croft are reproduced below:

$$D = 2r_0 \left\{ (\gamma - \beta_0) - \frac{r_0 \cos \beta_0}{2\sqrt{C}} \ln \frac{B^2 - 4AC}{4C \left( \sin \gamma + \frac{1}{r_b} \sqrt{C} + \frac{1}{2\sqrt{C}} B \right)^2} \right\} \quad (3.2a)$$

$$P' = 2 \left\{ r_b \sin \gamma - r_0 \sin \beta_0 + \frac{1}{A} \left[ -r_b \sin \gamma - \frac{B}{4\sqrt{A}} \ln \frac{B^2 - 4AC}{(2Ar_b + B + 2r_b \sqrt{A} \sin \gamma)^2} \right] \right\} \quad (3.2b)$$

$$P = 2 \left\{ -r_0 \sin \beta_0 + \frac{B}{4} \left[ \frac{1}{\sqrt{A}} \ln \frac{B^2 - 4AC}{4 \left( Ar_b + \frac{B}{2} + \sqrt{Ar_b} \sin \gamma \right)^2} + \frac{r_m}{\sqrt{C}} \ln \frac{B^2 - 4AC}{4C \left( \sin \gamma + \frac{1}{r_b} \sqrt{C} + \frac{1}{2\sqrt{C}} B \right)^2} \right] \right\} \quad (3.2c)$$



Graph 3.1 Power loss as a function of polarisation angle.

where

$$F = \frac{f}{f_c}$$

$$A = 1 - \frac{1}{F^2} + \left( \frac{r_b}{F y_m} \right)^2 \quad (3.3a)$$

$$B = -\frac{2r_m r_b^2}{F^2 y_m^2} \quad (3.3b)$$

$$C = \left( \frac{r_b r_m}{F y_m} \right)^2 - r_0^2 \cos^2 \beta_0 \quad (3.3c)$$

These equations were derived for the no-field case, so an effective frequency shift based upon the Quasi-Longitudinal approximation was used to determine the ray parameters for the ordinary and extraordinary waves. Given these values, the total polarisation rotation between two end points can be calculated as:

$$\varphi = \frac{1}{2} (\varphi^+ - \varphi^-) \quad (3.4a)$$

$$\therefore \varphi^\pm = \frac{2\pi f}{c} P^\pm \quad (3.4b)$$

where  $P^+$  is the phase path of the ordinary wave and  $P^-$  the phase path of the extraordinary wave evaluated for the desired ground range and the frequencies:

$$f^\pm = f_0 \pm f_{gm} \quad (3.5)$$

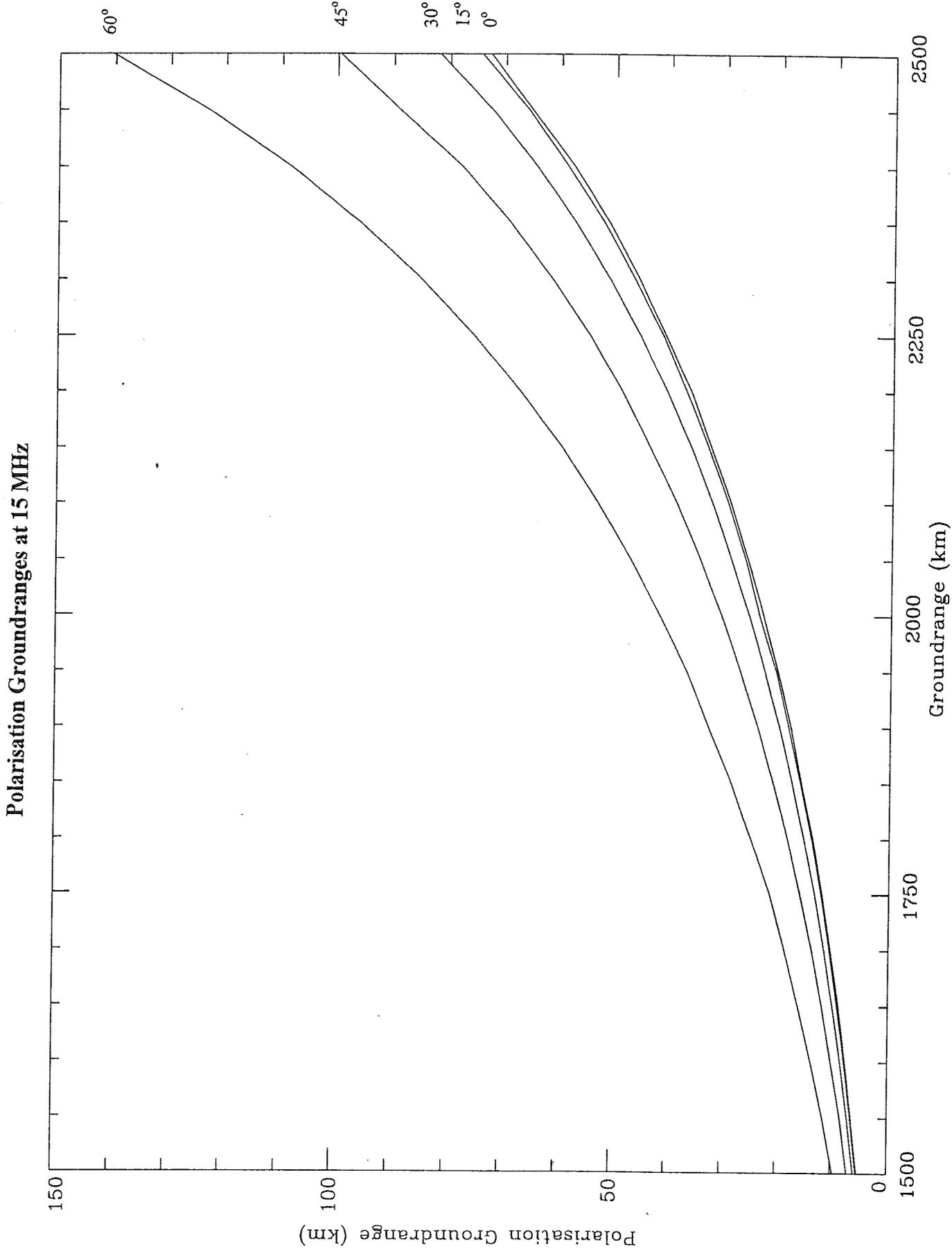
The initial elevation angles  $\beta_0^\pm$  were evaluated for a given ground range  $D_0$ , and an operating frequency  $f_0$  by solving equation (3.2a) numerically and substituting this value into (3.2b) and (3.2c).

### 3.2 SPATIAL DIVERSITY

The change in polarisation angle with ground range was investigated to determine what variation in ground range will cause the polarisation to go from horizontal to vertical for targets located within the radar's footprint.

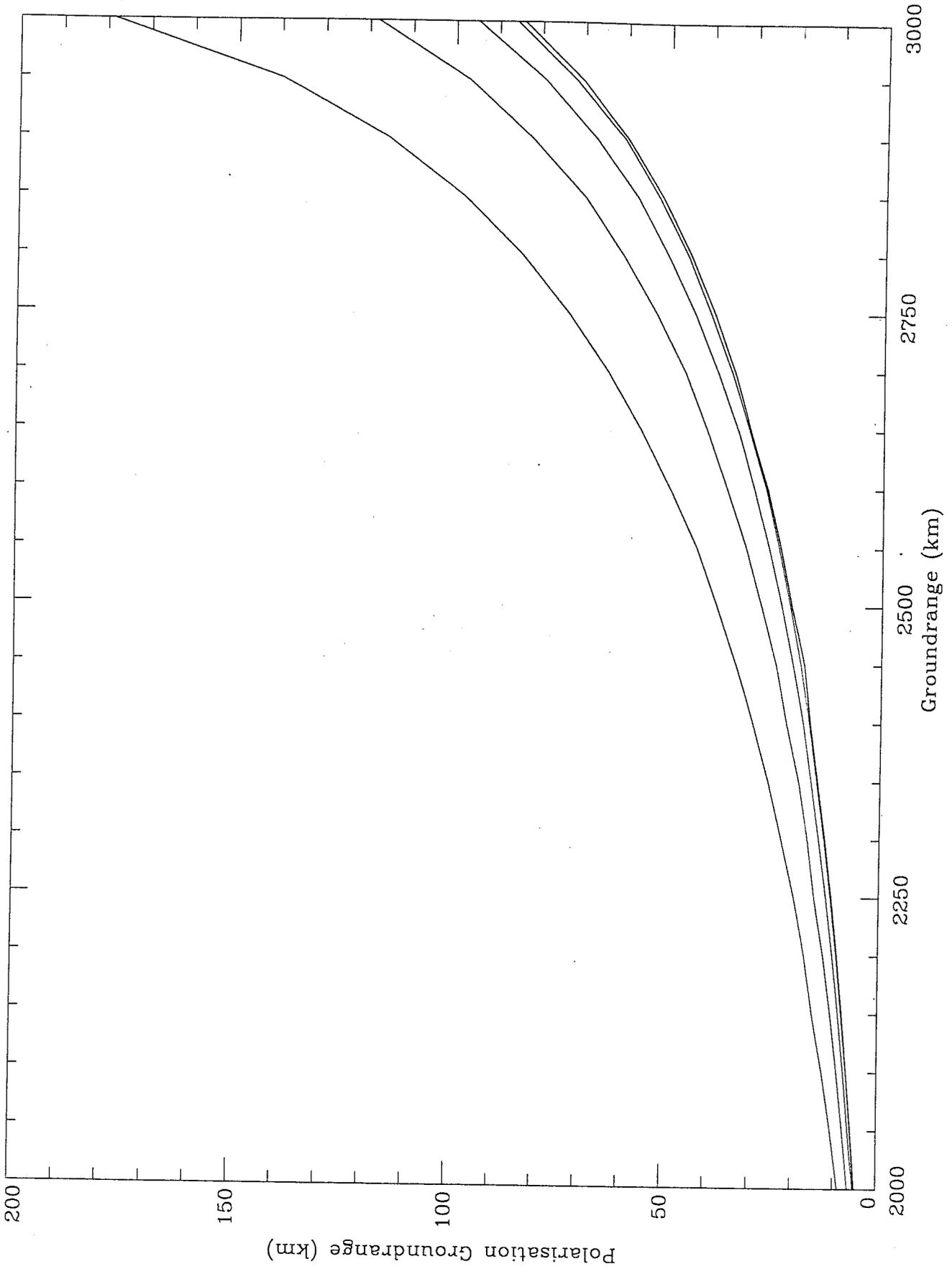
The required change in ground range (defined here to be the polarisation ground range) was evaluated by finding numerically what  $\Delta D$  corresponds to a  $\Delta\varphi$  of  $90^\circ$ . Using this technique the polarisation ground range was found to be in the tens to hundreds of kilometres and azimuthal variations for operational frequencies of 15MHz and 20MHz are shown in Graphs 3.2 and 3.3.

Thus given these ground ranges, one can see that even for an aircraft travelling at approximately 1000 kph, it would take between half a minute and ten minutes for the target's backscatter to change from being horizontally to vertically polarised. Given this slow rate of change of polarisation, any method that relies solely upon spatial diversity would require many dwells before a target hidden in a null would become visible, even if a rotation of only 30 degrees is required. The regions of illumination when the radar is operating at 80% of the MUF are shown in Graph 3.4. It should be noted that while these distances were calculated using the Quasi-Parabolic model, the introduction of even small-scale TIDs could cause

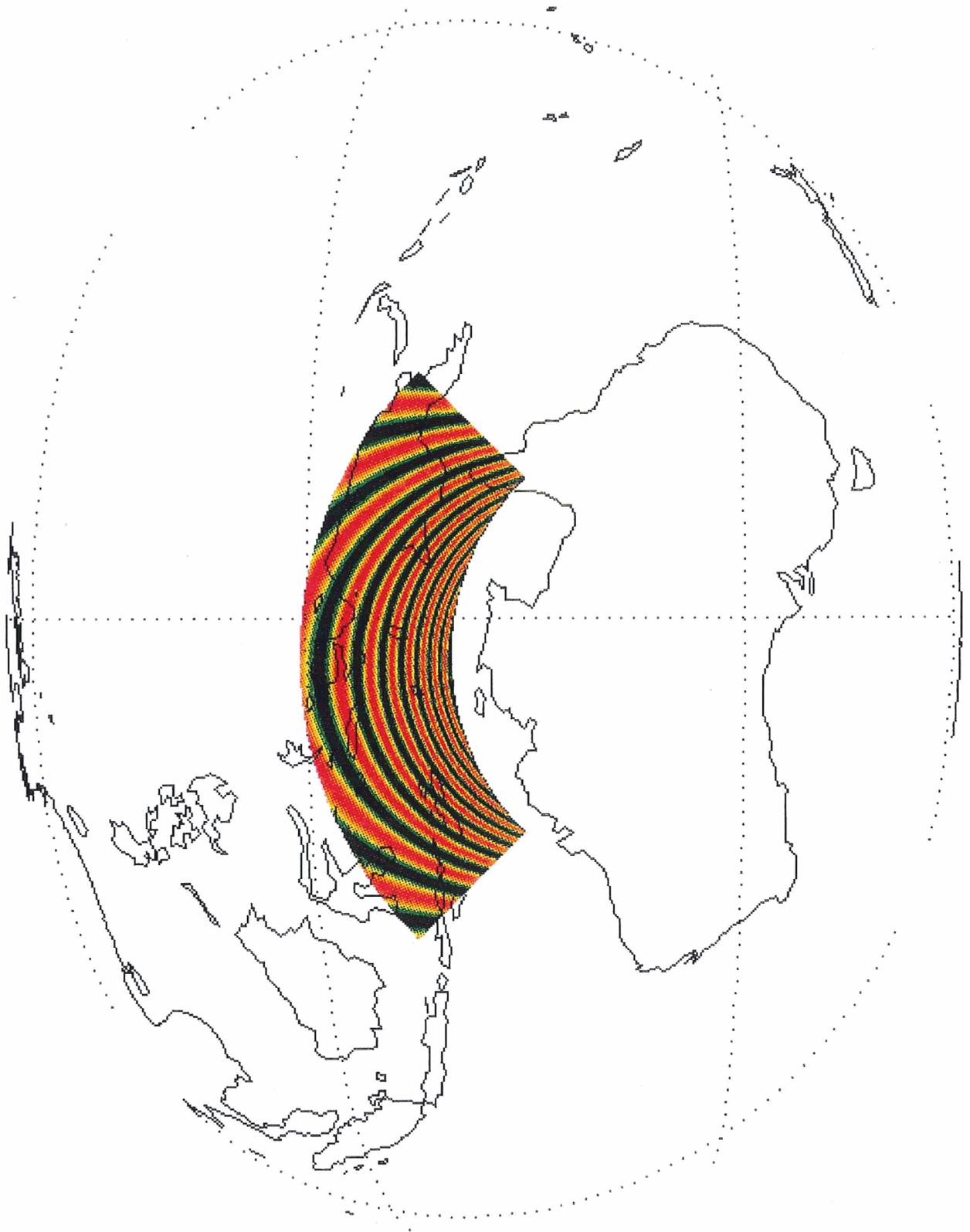


Graph 3.2 Polarisation Ground ranges for 15 MHz.

Polarisation Groundranges at 20 MHz



Graph 3.3 Polarisation Ground ranges for 20 MHz.



Graph 3.4 Geographic distribution of polarisation fringes.

significant deviations from the predicted ground ranges due to path differences associated with a non-stratified ionospheric model.

### 3.3 CHANNEL SEPARATION

The use of channel separation takes advantage of the dependence of polarisation upon the operating frequency. Thus, to achieve a desired change in the polarisation of the backscatter from a specific ground range, a corresponding frequency shift could be introduced.

In this report, two possible ways that channel separation could be used to reduce polarisation fading in OTHRs are examined; one involves operating the radar at two alternating frequencies such that the backscatter's polarisation changes by  $90^\circ$ , the other would involve measuring the polarisation of the backscatter from the desired ground range, and changing the operating frequency such that the backscatter incident at the receive site is now vertically polarised.

While the former is simpler to implement and effectively allows the reception of both horizontally and vertically polarised rays, because of the time multiplexing, each virtual array is useable only 50% of the time, In contrast, the frequency correction method would allow for optimum target detection by ensuring the received backscatter remains approximately vertically polarised. Alternatively, the real time polarisation correction could also be used to ensure that the transmitted signal is horizontally polarised when incident upon oceanic regions, thus maximising the signal to clutter ratio [5]. To do this, an approximation to  $\frac{\partial\phi}{\partial f}$  could be used to calculate the required frequency jump analytically.

One obvious problem in implementing this frequency-shift is that ionospheric perturbations such as fluctuations in electron density can cause predicted frequency changes to quickly become inaccurate, while a non-stratified layer can cause significant path deviations due to the aperiodic modulation of the refractive index along the ray's path of propagation. Thus any method that utilises frequency correction would need to taken such variations into account.

To investigate the magnitude of the frequency shift required to cause a  $90^\circ$  rotation (polarisation bandwidth), equation (3.4a) was solved numerically to find what  $\Delta f$  corresponded to a  $\Delta\phi$  of  $90^\circ$  at any given operating frequency. Below are specific polarisation bandwidths as a function of azimuth, with additional values shown graphically in Graph 3.5 to 3.7.

Operating Frequency (MHz)	Polarisation Bandwidth (kHz)			
	0° E of N	15° E of N	30° E of N	45° E of N
10	220.9	224.6	251.3	290.0
12	164.1	167.0	181.9	216.3
14	120.4	122.3	133.3	158.7
16	84.9	86.8	94.9	112.4
18	56.0	57.1	62.1	74.6
20	30.6	31.1	34.3	40.9

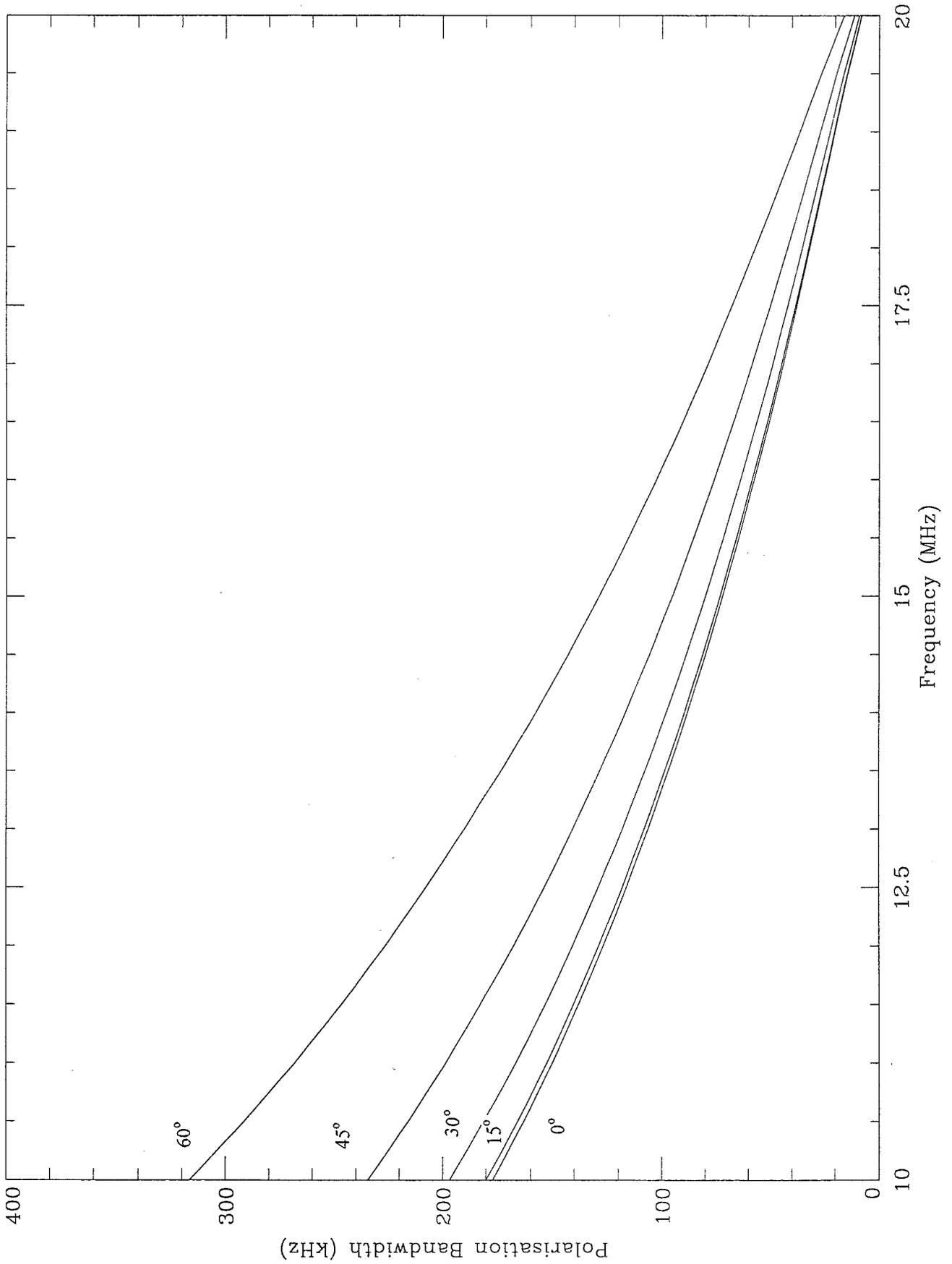
Table 3.1 Polarisation Bandwidths for selected frequencies and azimuths for a 2000km ray path.

These values agree with the calculations described in [4]. Given a fixed ground range, it was found that for changes in frequency of the order of the polarisation bandwidth,  $\frac{\partial\phi}{\partial f}$  is almost constant. Thus, one way to implement the frequency correction method would involve using a linearised approximation to the gradient which gives:

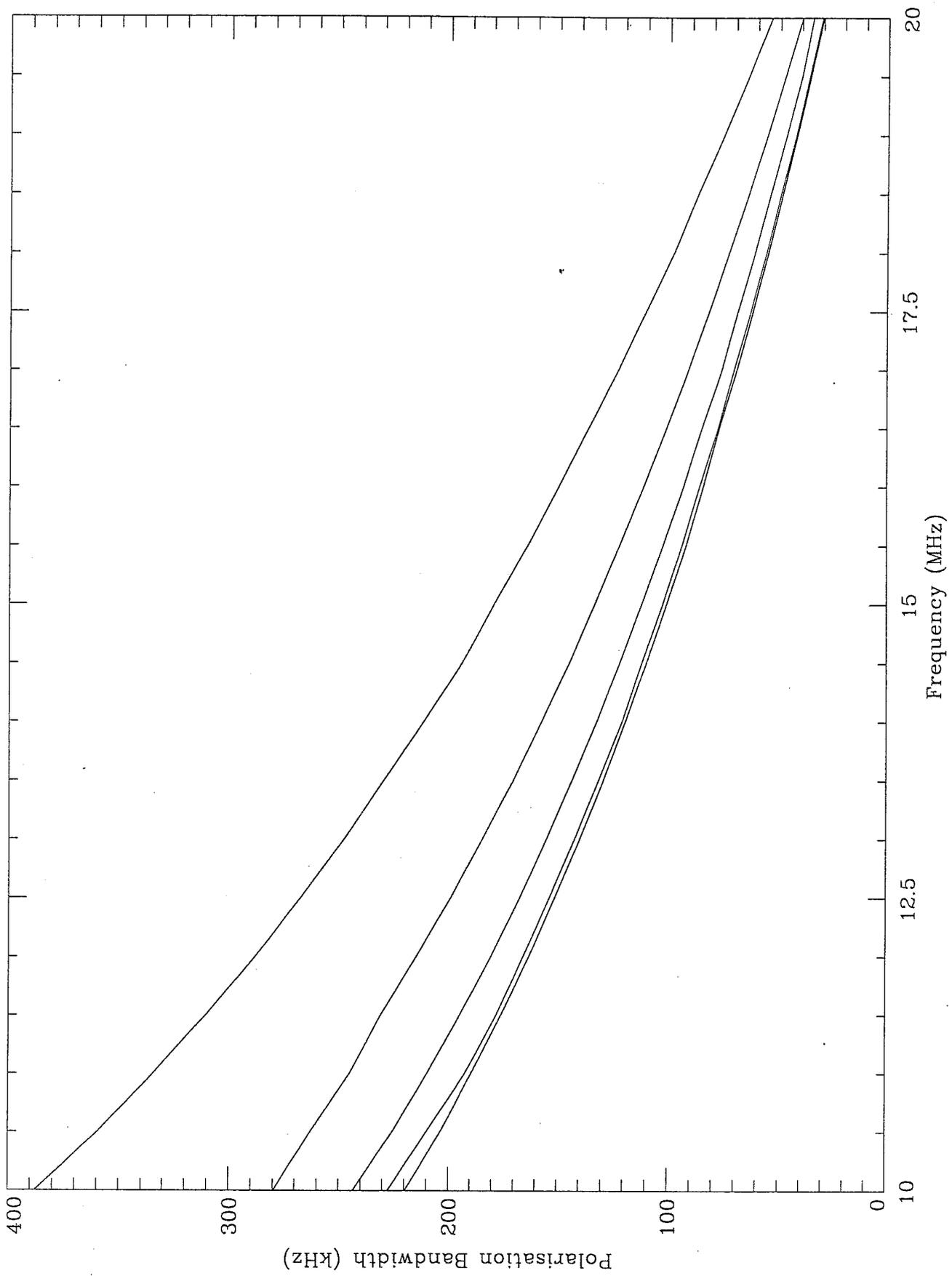
$$\left. \frac{\Delta f}{\Delta\phi} \right|_{f_0} = \frac{c}{\pi [(P'^+ - P^+) - (P'^- - P^-)]} \quad (3.6a)$$

$$\therefore \left. \frac{\partial\phi}{\partial f} \right|_{f_0} = \frac{\pi}{c} [(P'^+ - P^+) - (P'^- - P^-)] \quad (3.6b)$$

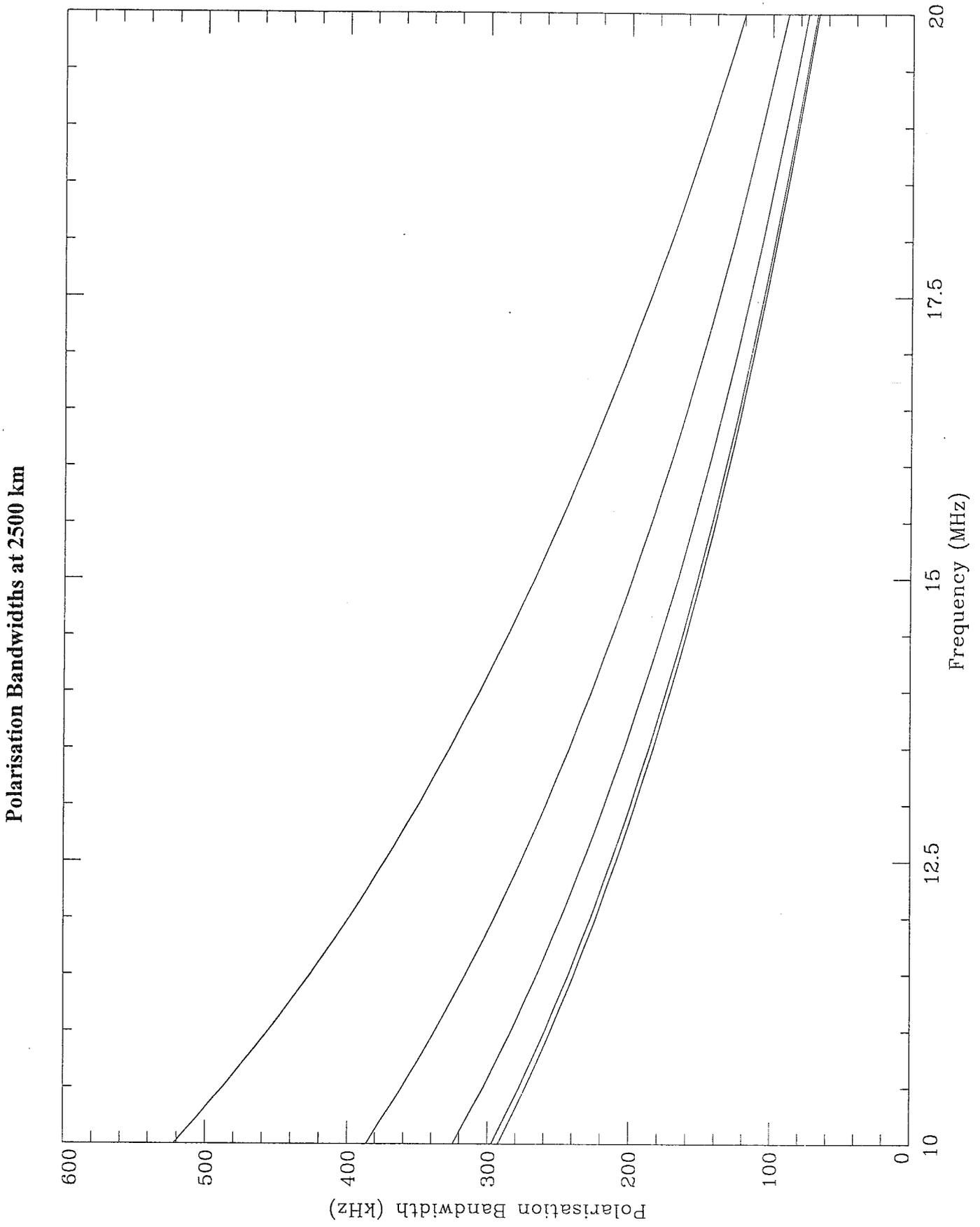
Polarisation Bandwidths at 1750 km



Graph 3.5 Polarisation Bandwidths for a ground range of 1750km.



Graph 3.6 Polarisation Bandwidths for a ground range of 2000km.



Graph 3.7 Polarisation Bandwidths for a ground range of 2500km.

where

$$\begin{aligned} \frac{\partial \phi^{\pm}}{\partial f} &= \frac{\partial \phi^{\pm}}{\partial P} \frac{\partial P}{\partial f} \\ &= \frac{2\pi f}{c} \times \frac{1}{f} (P'^{\pm} - P^{\pm}) \end{aligned} \quad (3.7)$$

It was found that the approximation (3.6a) accurately predicts the change in frequency required to rotate the polarisation vector to within a few degrees of  $\Delta\phi$ , as shown in Graph 3.8, where the error is the difference between the actual  $\Delta\phi$  and the calculated  $90^{\circ}$  rotation. One operational consequence is that under certain atmospheric conditions, it may be possible to estimate the required change in frequency based upon purely empirical data. Given this linear relationship, the gradient could then be evaluated by measuring the polarisation angle at two distinct operating frequencies, although temporal variations in the ionosphere's structure will affect the period of time for which this approximation remains valid.

### 3.4 COMPARISON WITH RAY TRACED VALUES

As stated earlier, all the results previously discussed were generated using the simple quasi-parabolic model of the ionosphere, and thus are subject to the limitations of the model. To investigate how reliable these relationships are when a dynamic ionospheric model is introduced, a raytracing package developed by Coleman[6] was used into which travelling ionospheric disturbances were introduced.

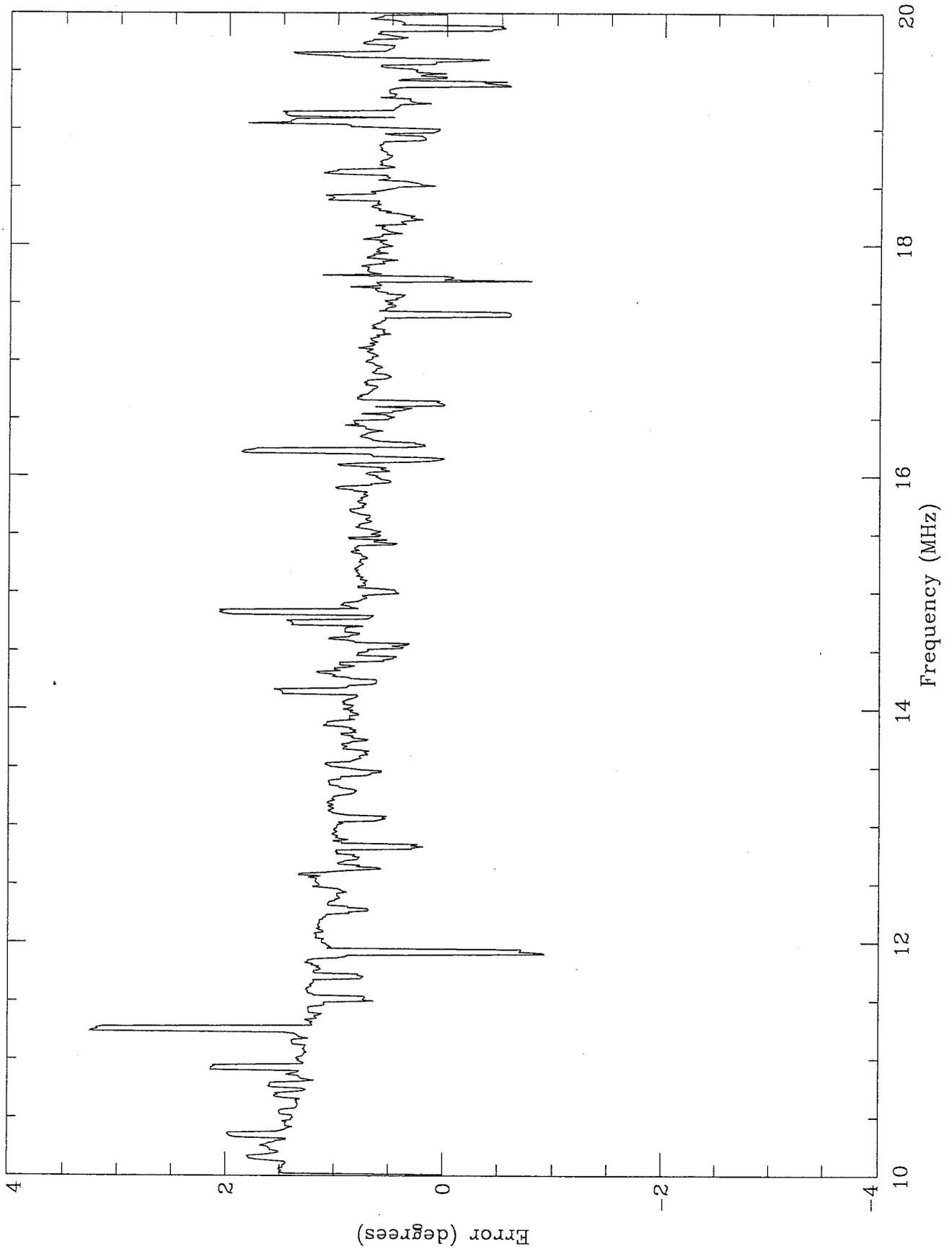
Under static conditions, the raytracing package agreed closely with the both the exact and approximate analytic results (see Table 3.2), but these calculations became less accurate as larger scale disturbances were introduced.

Operating Frequency (MHz)	Polarisation Bandwidth (kHz)		
	Croft's QP	Linearised Model	Ray Traced
10	220.9	224.7	230.9
12	164.1	166.1	161.9
14	120.4	121.6	118.9
16	84.9	85.9	78.6
18	56.0	56.3	52.2
20	30.6	31.0	30.7

Table 3.2 Comparison of Polarisation Bandwidths for a 2000km ray path, obtained using different models.

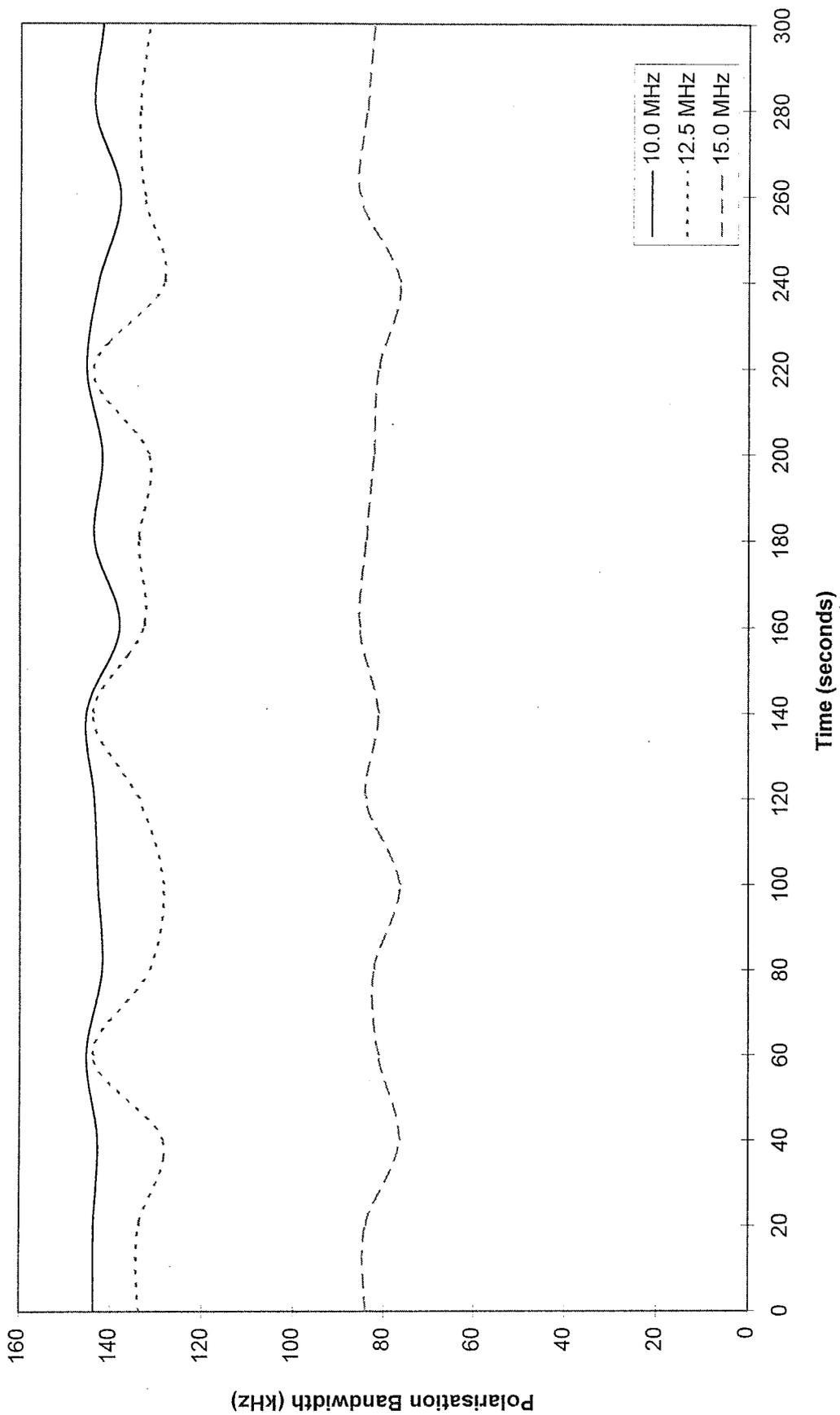
To determine how sensitive the linearised predictions are to changes in ionospheric parameters, a sinusoidal TID that modulated the maximum electron density was introduced into the system. As can be seen from Graphs 3.9 and 3.10, the introduction of the TID likewise results in a modulation of the polarisation bandwidth over time. When the TID's modulation was set to 10% variations of up to 20kHz were found, while at 40% the polarisation bandwidth deviated by as much as 40kHz from the homogeneous case.

Given the magnitude of these fluctuations, it is obvious that any method that uses  $\frac{\partial \phi}{\partial f}$  to predict a frequency jump can quickly become inaccurate once the maximum electron density, and thus the polarisation bandwidth varies by more than several percent. Using the initial gradient (ie. at  $t=0$ ), frequency corrections were made when ever the incoming ray approached  $60^{\circ}$  off vertical, and as can be seen from Graph 3.11, even with a TID modulation of only 5%, the changing electron density can cause a prediction to be as far as  $20^{\circ}$  off, although even this error corresponds to less than 1dB of loss. Thus, while the linear relationship can be used to accurately predict  $\Delta f$  given  $\Delta\phi$ , any technique that tries to actively minimise polarisation losses will need to take into account the perturbations caused by ionospheric irregularities and TIDs.



Graph 3.8 Error in predicting a 90° correction using the linearised model.

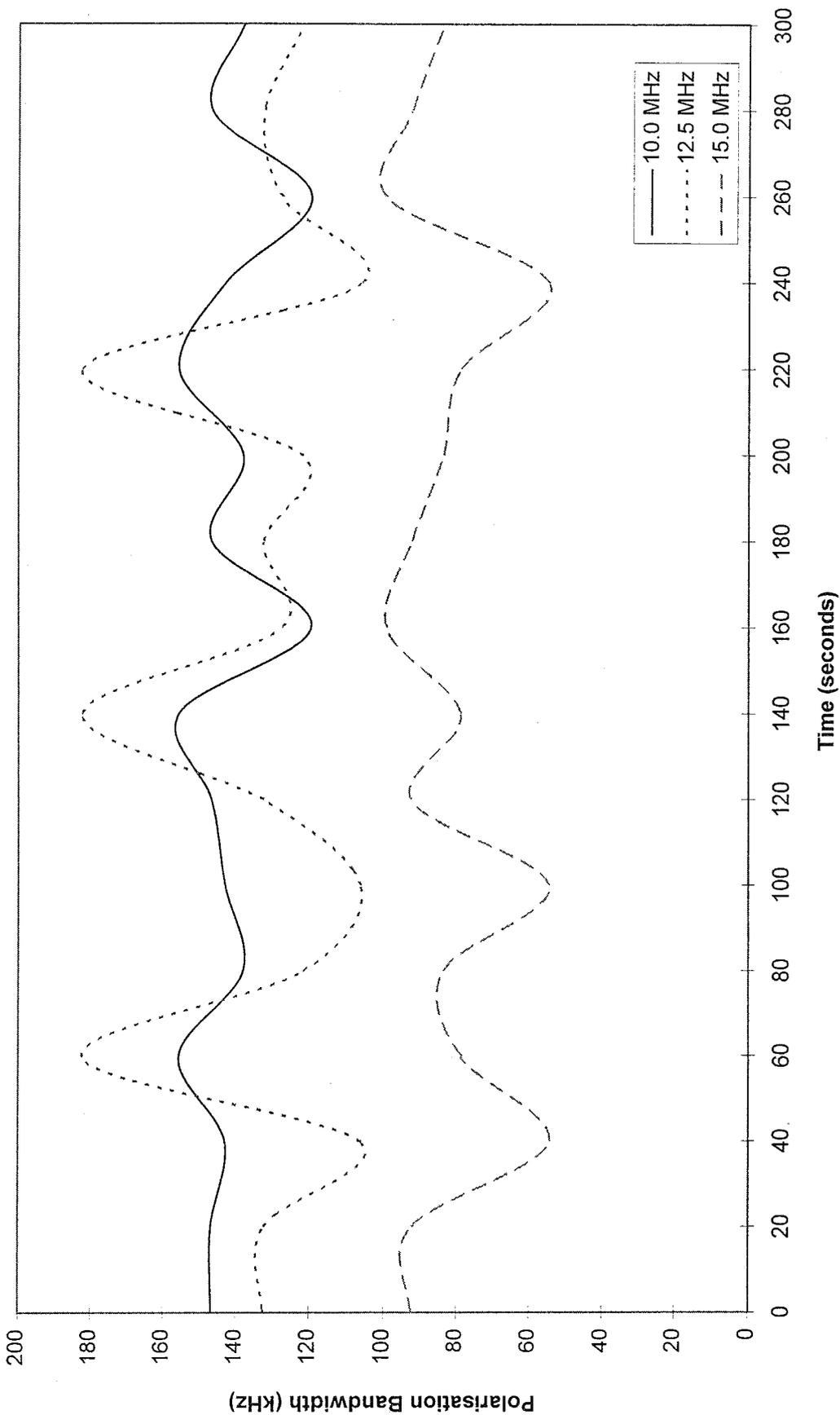
Effect of TID on Polarisation Bandwidth (modulation=10%)



Graph 3.9

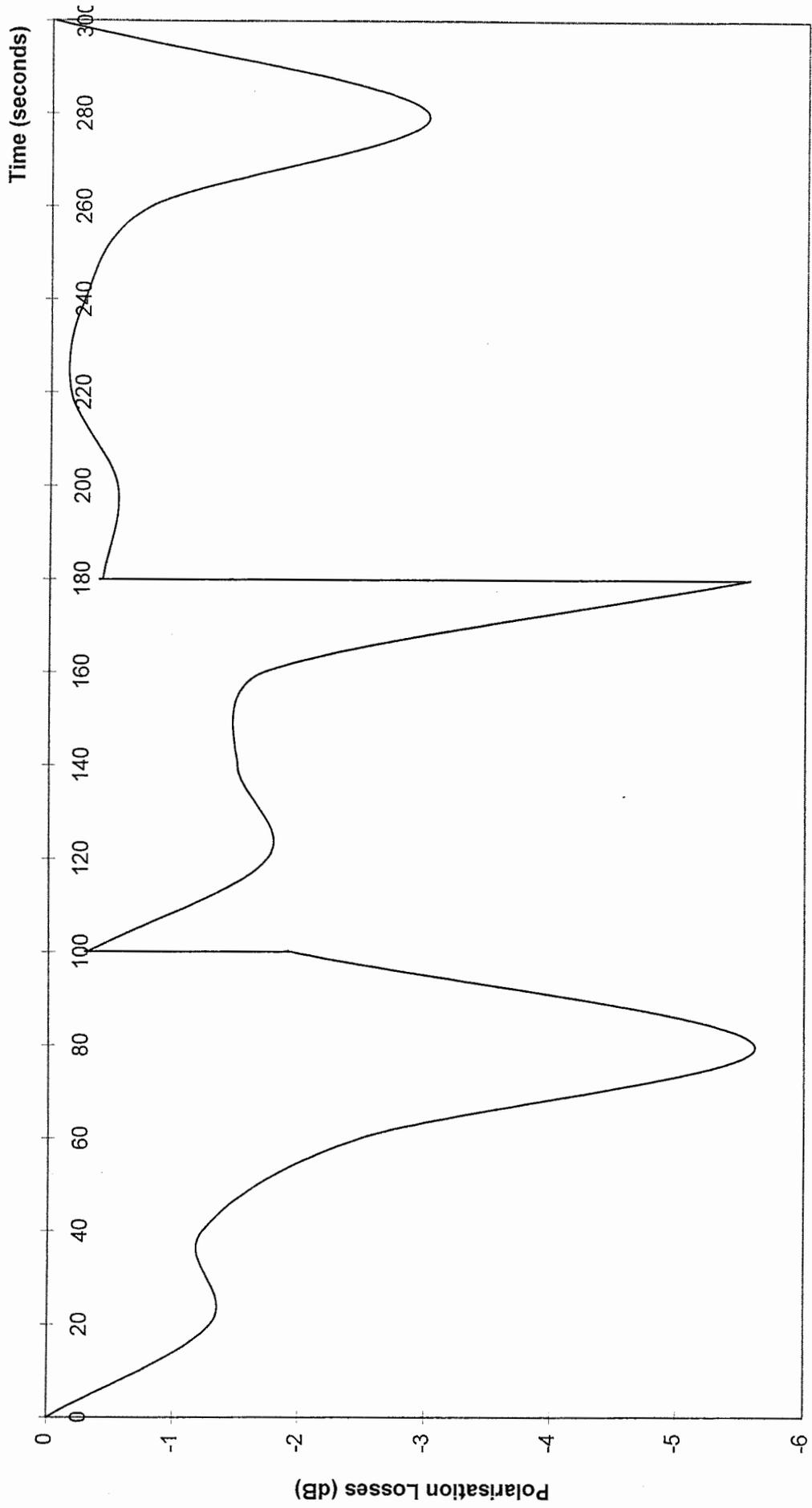
TID modulation of Polarisation Bandwidths.

Effect of TID on Polarisation Bandwidth (modulation=40%)



Graph 3.10 TID modulation of Polarisation Bandwidths.

Polarisation Losses using Frequency Correction (TID modulation 5%)



Graph 3.11

Effect of frequency corrections on polarisation losses.

#### 4. OTHER CAUSES OF FADING

While the focus of this project has been the effects of polarisation fading on HF radar systems, it is by no means the only cause of signal fading; there are other phenomena that can cause significant fluctuations in the backscatter's amplitude, of which interference and diffraction fading are the most significant. Apart from being able to resolve reflections off different ionospheric layers and thus eliminate the multimodal fading that plagues other HF systems, OTH radars are also relatively immune to absorption and other forms of fading whose period is much greater than the dwell time. Shown below are the most common types of HF signal fading, and their approximate periods [7, 8]:

Fading Type	Cause	Fading Period	Relative fading depths
Flutter	Small irregularities in the F-layer	10-100 ms	medium
Diffraction	Movement of ionospheric irregularities	5-20 s	medium
Polarisation	Rotation of polarisation ellipse	10-100 s	large
Interference (focusing)	Changes in curvature of the reflecting layer	15-30 min	large
Absorption	Time variation of ionospheric absorption	60 min	small

Table 4.1 Summary of HF fading phenomena

Flutter is primarily caused by the high speed convection of irregularities in the F-region, where velocities above  $100\text{ms}^{-1}$  are not uncommon. Due to its short fading period, the effects of flutter should be visible in the raw received radar signal. Flutter has been found to have a fading correlation bandwidth of greater than 1kHz, and its distortion of AM signals is more severe than other modulated signals, due to the destructive interference caused by multiple rays being reflected off these F-layer irregularities. It is most common at times of spread-F, near sunset and especially near the equinoxes.

Diffraction fading occurs when a strong ionospheric irregularity causes destructive interference between waves that have diffracted around it. This phenomena, which usually lasts for less than twenty seconds but can persist for up to several minutes, can be modelled in terms of diffraction around a cylindrical non-absorbing irregularity. This form of fading is also frequency dependent, with the fading depths increasing proportionally with wavelength.

Interference fading is one of the most common forms of fading and is caused by interactions between multiple waves reflected off an irregular ionospheric layer. This phenomena is also described as focusing/defocusing and may occur on both oblique and vertical propagation paths, although its effects become less significant with increasing frequency. These perturbations in ionospheric reflectors (ie. any surface of constant electron density) have several causes including the passage of acoustic-gravity waves through the neutral ionosphere and large scale TIDs (up to thousands of kilometres long) travelling at speeds between  $100\text{ms}^{-1}$  and  $1000\text{ms}^{-1}$ .

As HF waves experience most of their signal attenuation while travelling through the D-layer, absorption fading is caused by changes in the structure of the D-region. Although it only exists during day-time, the movement of irregularities in the D region, changes in ionising solar radiation and monthly variations in absorption can cause noticeable fade outs. As these periods can last anywhere between minutes to around an hour, it is only discernible in OTH radars by long term variations in backscatter and clutter signal strengths, although the SCV should remain relatively unchanged.

## 5. REFERENCES

- [1] Earl, Dr G.F. (private communication), High Frequency Radar Division, Defence Science and Technology Organisation Salisbury
- [2] Croft, T.A. and Hoogasian, H (1968), "Exact Ray Calculations in a Quasi-Parabolic Ionosphere with no Magnetic Field", Radio Science Vol. 3 No. 1, January 1968
- [3] Davies, K. (1989), Ionospheric Radio (Peter Peregrinus Ltd, UK), chapters 6, 8
- [4] Epstein, M.R. (1967), "Computer Prediction of the Effects of HF Oblique-Path Polarisation Rotation with Frequency", Radioscience Laboratory, Stanford Electronics Laboratories
- [5] Barnum, J.R. (1970), "The Effect of Polarisation Rotation on the Amplitude of Ionospherically Propagated Sea Backscatter", Radioscience Laboratory, Stanford Electronics Laboratories
- [6] Coleman, Dr C. (private communication), High Frequency Radar Division, Defence Science and Technology Organisation Salisbury
- [7] Maslin, N. (1987), HF Communications: A System Approach (Pitman Publishing, London), p77
- [8] Davies, K. (1989), Ionospheric Radio (Peter Peregrinus Ltd, UK), chapters 7

## APPENDIX A: Derivations and Proofs

$$\begin{aligned}
 E_{B,V} &= E \cos\theta & E_{B,H} &= \beta_{BE} E \sin\theta \\
 E_{R,V} &= E \cos\theta \cos\phi & E_{R,H} &= -\beta_B^2 E \sin\theta \cos(\phi + \frac{\pi}{2}) \\
 & & &= \beta_B^2 E \sin\theta \sin\phi \\
 \Rightarrow E_R &= E(\cos\theta \cos\phi + \beta_B^2 \sin\theta \sin\phi)
 \end{aligned}$$

**Equation A.1 Induced EMF at the receive site using beacon method 2.**

$$\begin{aligned}
 E_{B,V} &= \frac{E}{2} [\cos\theta - \beta \sin\theta] & E_{B,H} &= \frac{\beta_B E}{2} [\beta_B \sin\theta - \cos\theta] \\
 E_{R,V} &= \frac{E}{2} [\cos\theta \cos\phi + \beta_B^2 \sin\theta \sin\phi] & E_{R,H} &= \frac{\beta_B E \sin\phi}{2} [\beta_B \sin\theta - \cos\theta] \\
 \Rightarrow E_R &= \frac{E}{2} [\cos\theta \cos\phi - \beta_B \sin\theta \cos\phi + \beta_B^2 \sin\theta \sin\phi - \beta_B \cos\theta \sin\phi]
 \end{aligned}$$

**Equation A.2 Induced EMF at the receive site using beacon method 3.**

$$\begin{aligned}
 E_{B,V} &= E \cos\theta & E_{B,H} &= \beta_B^2 E \sin\theta \\
 E_{R,V} &= E[\cos\theta \cos\phi + \beta_B^2 \sin\theta \sin\phi] & E_{R,H} &= \beta_R E[\beta_B^2 \sin\theta \cos\phi - \cos\theta \sin\phi] \\
 \Rightarrow E_R^2 &= E^2 [\cos\theta(\cos\phi + \beta_R \sin\theta) + \beta_B^2 \sin\theta(\sin\phi + \beta_R \cos\phi)]^2
 \end{aligned}$$

**Equation A.3 Induced EMF at the receive site after introducing the horizontal antenna array**

Define  $\beta_B \in [0,1]$ , since we assume finite antenna gain.

Then,  $P_R = \|E_R\|^2$

$$\begin{aligned}
 \frac{\partial P_R}{\partial \beta_B} &= \frac{\partial}{\partial \beta_B} \|E_R\|^2 \\
 &= 2\beta_B \|\sin\theta \sin\phi\| \times 2 \|\cos\theta \cos\phi + \beta_B^2 \sin\theta \sin\phi\| \\
 &= 0 \quad \forall \theta, \phi \Leftrightarrow \beta_B = 0
 \end{aligned}$$

but,

$$\begin{aligned}
 \frac{\partial^2 P_R}{\partial \beta_B^2} &\geq 0 \\
 \Rightarrow \beta_B &= 1 \text{ maximises received power.}
 \end{aligned}$$

**Equation A.4 Proof for optimal beacon gain ratio**

## APPENDIX B: List of Symbols

Symbol	Description
$\theta$	Polarisation rotation on outbound (transmitter-beacon) leg
$\phi$	Polarisation rotation on inbound (beacon-receiver) leg
$\gamma$	Angle ray makes at the bottom of the ionosphere
$\varphi$	Polarisation (Faraday) rotation in radians
$\beta_0$	Initial elevation angle of ray
$\beta_B$	Voltage gain ratio ( $G_H:G_V$ ) of beacon's antennas
$\beta_R$	Voltage gain ratio of receiver's antennas
$D$	Ground range (metres)
$E$	Amplitude of transmitted EMF
$E_R$	Instantaneous received EMF
$N_M$	Maximum electron density ( $10^{12} \text{ m}^{-3}$ )
$P_R$	Instantaneous receive power
$P$	Phase path (metres)
$P'$	Group path (metres)
$r_0$	Radius of the earth (6370km)
$r_b$	Radius to the layer base ( $r_m - y_m$ )
$r_m$	Radius at which electron density is at the maximum (300km)
$y_m$	Layer semithickness (100km)

## APPENDIX C: Source Code and Explanation of Programs

Provided here is a brief description of the major programs used to generate the graphs and probability distributions detailed in this document followed by the source code itself:

### ***RECEIVE\_PLOT***

Used to generate Graphs 2.1 to 2.4, and Graphs 2.8 and 2.9. Calculates received power as a function of  $\phi$  and  $\theta$  (polarisation angles on outgoing and incoming legs) and  $\beta_B$  for a user specified  $\beta_R$ , and then collates these signal powers to derive the probability distributions shown in Graphs 2.8 and 2.9. These are calculated by finding the ratio of the number ( $\phi, \theta$ ) pairs whose return power is above a user specified level to the total number of points for a given  $\beta_B$ .

### ***BEACON\_PLOT***

Used to generate Graphs 2.5 through 2.7. Displays the voltage induced at the beacon as a function of polarisation angle and antenna gain ratios. Calculates the data for  $\beta_B = 0$  through to the user specified maximum (MAX), and duplicates it for  $\beta_B = -MAX$  to  $\beta_B = 0$ .

### ***POLARISATION***

Used to generate Graphs 3.2 through 3.8. Implementation the equations described in chapter 3 to calculate the polarisation rotation over a given ground range, the polarisation ground ranges and bandwidths as well as their linearised approximations. Also included are some routines that were used to evaluate the linearised model's accuracy.

### ***ROUTINES***

These routines are required by all other programs, and include the functions to solve the initial elevation angles, effective frequency shifts as well as calculate ray parameters such as ground range, phase path and group paths. The routine is courtesy of Dr. Chris Coleman.

RES\$IEGROUP: [IDL.LIBRARY. DREISIGER.RESEARCH.OLD] BEACON\_PLOT.PRO;5

;author: Peter Dreisiger, Dec 1996 IE HFRD DSTO SALISBURY

;This programme shows how the reradiated beacon power varies with BETA (in dB)  
;using an uncoupled horizontal-vertical antenna pair

FUNCTION Mirror, dat

```

fin = findgen(720, 360)
ind = 359 - indgen(360)
fin(0:359, * ) = dat
fin(360: * , * ) = dat(ind, * )
fin = congrid(fin, 360, 360)
RETURN, fin

```

END

PRO Plot\_graph, dat, title

```

erase
x_beta_value = ['-3', '-2', '-1', '0', '1', '2', '3']
y_theta_value = [' 0', ' 45', ' 90', '135', '180', '225', '270', '305', '360'
]
a = [1, 2]

```

```

plot, a, /NODATA, POS=[98, 98, 460, 460], /DEVICE, TICKLEN= -0.02, XTICKS=6, $
YTICKS=8, XMINOR=5, YMINOR=9, XTICKNAME=x_beta_value, $
YTICKNAME=y_theta_value, XTITLE='!6Beta (dB)', YTITLE='!6Theta (degrees)'
max_val = max(dat, MIN=min_val)
dat = ie bytscl(dat, DMIN=min_val, DMAX=max_val, TOP=!colour.top)
color_scale, /VERT, MINVAL=100 * min_val, INT=10, MAJ=ceil((max_val - min_val)
* 10), $

```

```

XPOS=500, YPOS=100, BAR_LENGTH=360, TITLE='Percentage Retransmitted ' + $
'Power', MAX_USEABLE_COL=!colour.top
xyouts, 30, 520, title, CHAR_SIZE=2, /device
tv, dat, 100, 100, /DEVICE

```

END

PRO Beacon\_plot

```

window, TITLE='Passive Beacon Return Power Analysis', XSIZE=800, YSIZE=600
load_colours, "FMS Rainbow"
beta = 6.0
read, beta, PROMPT='Enter the maximum absolute value for Beta (db): '
beta_max = exp( -alog(10) /20 * beta)

```

```

theta_arr = findgen(360) /57.3
beta_arr = beta_max + (findgen(360) /360 * (1 - beta_max) )
identity = fltarr(360) +1
beta_arr = beta_arr^2
pby = cos(theta_arr)^2
pby = mirror(pby##identity)
pbx = sin(theta_arr)^2
pbx = mirror(pbx##beta_arr)
tot = pbx + pby

```

```

plot_graph, pbx, '!6Power Received by Horizontal Beacon Antenna'
hak, /MESSAGE
plot_graph, pby, '!6Power Received by Vertical Beacon Antenna'
hak, /MESSAGE
plot_graph, tot, '!6 Total Received Power at Beacon'

```

END

```

;author: Peter Dreisiger, Dec 1996 IE HFRD DSTO SALISBURY
; This program generates power v's polarisation plots across beta (beacon) and
; beta (receiver) space. It also allows the user to analyse the probability of
; detecting signals with Xdb of the maximum detectable signal, and can view
; this data as either plots or animations
;=====
PRO Receive_plot
clear
load colours, 'FMS Rainbow' ;define constants
theta = findgen(360) /57.3
cs = cos(theta)
sn = sin(theta)
vx = cs**ics
vy = sn**sn
hx = sn**ics
hy = cs**sn
a = [1, 2]
lab = [' 0', ' 45', ' 90', '135', '180', '225', '270', '305', '360']
thresh = 0.0
steps = 3.0
bet = 3.0

;=====
; obtain parameters
print, & print & print & print & print
PRINT, '=====
read, bet, PROMPT='Please specify the maximum absolute value of BETA (dB): '
read, steps, PROMPT='Enter the number of increments BETA will go through: '
read, beta, PROMPT='Enter the magnitude of BETA for the receiver in dB: '
PRINT, '=====
print
print, 'Please wait... Allocating memory'
print
bet = exp(-alog(10)/10 * bet)
rbeta = 10*( -rbeta /10)
power = filtarr(steps+1, 360, 360)
prob = filtarr(steps+1, 2)
window, TITLE='Passive Beacon Return Signal Analysis'

;generate data slices
print, 'Please wait... Generating data'
print
FOR BETA=0, steps DO BEGIN
print_bar, beta, steps
sbeta = (bet +float(beta) * (1.0 -bet) ) /steps)^2
power(beta, *, *) = 10 * alog10((vx +sbeta*vy)^2 $
+ (rbeta*(sbeta * hx -hy))^2)
ENDFOR

choice = ''
WHILE (choice NE '-1') DO BEGIN
print & print & print & print
PRINT, '=====
print, 'Press 1 to view (theta,phi) at a specific value of Beta'
print, ' 2 to view the probability distributions for a specific S-N ra
tio'
print, ' 3 to view all (theta,phi) slides in an animation'
print, ' -1 to quit'
PRINT, '=====
read, choice, PROMPT='Please enter your choice: '

CASE choice OF
/1: BEGIN
;view (theta,phi) plots

```

```

beta = 0.0
read, beta, PROMPT='Enter a value for Beta in dB (-1 to quit)'
WHILE (beta GE 0) DO BEGIN
beta = fix(steps / (1.0 -bet))*( -bet +10.0*( -beta /10.0))
IF (beta GT steps) THEN BETA=steps
IF (beta LT 0) THEN beta = 0
actb = -10 * alog10(bet +float(beta) * (1.0 -bet) ) /steps
print, 'Retrieving slide for BETA=' +string(actb, FORMAT='f4.2')
plot, a, /NODATA, POS=[98, 98, 460, 460], /DEVICE, TICKLEN=-0.02, $
XTICKS=8, YTIKCS=8, XMINOR=9, YMINOR=9, XTICKNAME=lab, $
YTIKNAME=lab, XTITLE='!6Theta (degrees)', YTITLE='!6Phi (degrees)',
$
TITLE='Power Distribution for Beta = ' +string(actb, FORMAT='f4.2')
color scale, /VERT, MINVAL=-20, INT=5, MAJ=4, XPOS=500, YPOS=100, $
BAR LENGTH=360, TITLE='S/N (dB)', MAX_USEABLE_COL=:colour.top
tv, !e_bytscl(power(beta, *, *) , DMAX=0, DMIN=-20, $
TOP=:colour.top), 100, 100, /device

beta = 0.0
read, beta, PROMPT='Enter a value for Beta in dB (-1 to quit)'
ENDWHILE

/2: BEGIN
;collating and displaying statistics
read, thresh, PROMPT='Enter the cut-off ratio in dB: '
print, 'Please wait... Collating statistics'
print

FOR BETA=0, steps DO BEGIN
index = where(power(beta, *, *) GT -thresh, count)
prob(beta, 0) = -10 * alog10(bet +float(beta) * (1.0 -bet) ) /steps
prob(beta, 1) = float(count) /129600
print_bar, beta, steps
ENDFOR

erase
baxis = -10 * alog10(bet + (1.0 -bet) * findgen(101) /100.0)
plot, baxis, spline( -prob( *, 0), prob( *, 1), -baxis), XTITLE='!6Beta
' (dB)', YTITLE='!6Probability', POS=[98, 98, 560, 460], /DEVICE
read, thresh, PROMPT='Enter the cut-off ratio in dB (-1 to quit): '
WHILE (thresh GE 0) DO BEGIN
print, 'Please wait... Collating statistics'
print
FOR BETA=0, steps DO BEGIN
index = where(power(beta, *, *) GT -thresh, count)
prob(beta, 0) = -10 * alog10(bet +float(beta) * (1.0 -bet) ) /steps
prob(beta, 1) = float(count) /129600
print_bar, beta, steps
ENDFOR

baxis = -10 * alog10(bet + (1.0 -bet) * findgen(101) /100.0)
oplot, baxis, spline( -prob( *, 0), prob( *, 1), -baxis)
read, thresh, PROMPT='Enter the cut-off ratio in dB (-1 to quit): '
ENDWHILE

/3: BEGIN
;animate (theta,phi) plots
plot, a, /NODATA, POS=[98, 98, 460, 460], /DEVICE, TICKLEN=-0.02, $
XTICKS=8, YTIKCS=8, XMINOR=9, YMINOR=9, XTICKNAME=lab, $
YTIKNAME=lab, XTITLE='!6Theta (degrees)', YTITLE='!6Phi ' + $
' (degrees)'
color scale, /VERT, MINVAL=-20, INT=5, MAJ=4, XPOS=500, YPOS=100, $
BAR LENGTH=360, TITLE='S/N (dB)', MAX_USEABLE_COL=:colour.top
FOR BETA=0, steps DO BEGIN
tv, !e_bytscl(power(beta, *, *) , DMAX=0, DMIN=-20, $
TOP=:colour.top), 100, 100, /device

```

20-FEB-1997 11:19  
RES\$EGROF: [IDL.LIBRARY.\_DREISIGER.RESEARCH.OLD]RECEIVE\_PLOT.PRO;217

Page 3

DREISIGER

ENDFOR  
END

Else:  
ENDCASE  
ENDWHILE  
END

RES\$IEGROUPE:[IDL.LIBRARY.\_DREISIGER.RESEARCH]POLARISATION.PRO;1

```
; author: Peter Dreisiger, Jan 1997
; Plots dP/df assuming a quasi-parabolic concentric ionosphere
; using effective frequency to take into account the magnetic
; fields
```

```
FUNCTION PolarRotation, freq, dist, bearing, fc
```

```
; author: Peter Dreisiger, HFRD Jan 1997
; Calculates polarisation rotation of a radio wave as it travels
; from Alice Springs to the beacon DIST kms away.
```

```
=====
; input: freq = operating frequency (MHz)
;         dist = ground range (metres)
;         bearing = bearing (degrees east of north)
;         fc = critical frequency (MHz)
; output: return = polarisation angle (rads)
=====
```

```
freq = double(freq)
dist = double(dist)
bearing = double(bearing)
fc = double(fc)
dip = 0d0
decl = 0d0
fh = 0d0
po = 0d0
px = 0d0
dior = atan(1d0) /45d0
```

```
magnetic, 132.25, -17.9, dip, decl, fh
fgm = double(0.5 * fh * cos(dip * dior) * cos((decl - bearing) * dior) )
fRo = double((freq +fgm) /fc)
fRx = double((freq -fgm) /fc)
b0o = findb0(fRo, dist, findbmax(fRo) )
b0x = findb0(fRx, dist, findbmax(fRx) )
IF ((b0o NE -1) AND (b0x NE -1) ) THEN BEGIN
  po = double(phasepath(fRo, b0o) )
  px = double(phasepath(fRx, b0x) )
ENDIF
RETURN, atan(0d0, -1d0) * (po -px) * freq/300d0
END
```

```
FUNCTION FinddF, freq, fc, dist, bearing
```

```
; author: Peter Dreisiger, HFRD Jan 1997
; Calculates the change in operating frequency required to achieve
; a 90 degree rotation in polarisation angle at the receive site.
```

```
=====
; input: freq = operating frequency (MHz)
;         fc = critical frequency (MHz)
;         dist = ground range (metres)
;         bearing = bearing in degrees
; output: return = polarisation bandwidth (MHz)
=====
```

```
fmin = double(freq)
fmax = fmin +2.5
con = atan(1d0, 0d0)
accuracy = 0.0001
```

```
phase = double(polarrotation(fmin, dist, bearing, fc) )
pmin = double(0)
pmax = double(polarrotation(fmax, dist, bearing, fc) -phase)
diff = fmax -fmin
WHILE (diff GT accuracy) DO BEGIN
  pmid = double(polarrotation(0.5 *(fmin +fmax), dist, bearing, fc) -phase)
  IF ((con GE pmin) AND (con LT pmid) ) THEN fmax = 0.5 *(fmax +fmin)
  IF ((con GE pmid) AND (con LE pmax) ) THEN fmin = 0.5 *(fmax +fmin)
```

RES\$IEGROUPE:[IDL.LIBRARY.\_DREISIGER.RESEARCH]POLARISATION.PRO;1

```
pmin = double(polarrotation(fmin, dist, bearing, fc) -phase)
pmax = double(polarrotation(fmax, dist, bearing, fc) -phase)
diff = fmax -fmin
```

```
ENDWHILE
RETURN, (0.5 *(fmin +fmax) ) -freq
END
```

```
FUNCTION FinddD, dist, freq, fc, bearing
```

```
; author: Peter Dreisiger, HFRD Jan 1997
; Calculates the change in ground range required to achieve
; a 90 degree rotation in polarisation angle at the receive site.
```

```
=====
; input: dist = ground range (meters)
;         freq = operating frequency (MHz)
;         fc = critical frequency (MHz)
;         bearing = bearing in degrees
; output: return = polarisation groundrange (meters)
=====
```

```
dmin = double(dist)
freq = double(freq)
fc = double(fc)
dmax = dmin +500000
con = !Pi /2
accuracy = 0.005
pmin = double(polarrotation(freq, dmin, bearing, fc) )
con = pmin -!pi /2
pmax = double(polarrotation(freq, dmax, bearing, fc) )
diff = dmax -dmin
WHILE (diff GT accuracy) DO BEGIN
  pmid = double(polarrotation(freq, 0.5 *(dmin +dmax) , bearing, fc))
  IF ((con GE pmin) AND (con LT pmid) ) THEN dmin = 0.5 *(dmax +dmin)
  IF ((con GE pmid) AND (con LE pmax) ) THEN dmax = 0.5 *(dmax +dmin)
  pmin = double(polarrotation(freq, dmax, bearing, fc) )
  pmax = double(polarrotation(freq, dmin, bearing, fc) )
  diff = dmax -dmin
ENDWHILE
RETURN, (0.5 *(dmin +dmax) ) -dist
END
```

```
FUNCTION DpdF, freq, df, fc, dist
```

```
; author: Peter Dreisiger, HFRD Jan 1997
; Calculates the change in polarisation angle as a function of the
; change in operating frequency, using the linearised model.
```

```
=====
; input: freq = operating frequency (MHz)
;         df = change in frequency (MHz)
;         fc = critical frequency (MHz)
;         dist = ground range (meters)
=====
```

```
freq = double(freq)
df = double(df)
fc = double(fc)
dist = double(dist)
bearing = 0.0
dior = atan(1d0) /45d0
magnetic, 132.25, -17.9, dip, decl, fh
fgm = double(0.5 * fh * cos(dip * dior) * cos((decl - bearing) * dior) )

fo = freq +fgm
fx = freq -fgm
b0o = findb0(fo /fc, dist, findbmax(fo /fc) )
b0x = findb0(fx /fc, dist, findbmax(fx /fc) )
po = phasepath(fo /fc, b0o)
pdo = grouppath(fo /fc, b0o)
```

```
px = phasepath(fx /fc, b0x)
pdx = grouppath(fx /fc, b0x)
RETURN, df * atan(0d0, -1d0) /300d0 *(pdo +px -po -pdx)
END
```

```
FUNCTION DfdP, freq, fc, dp, dist, bearing
; author: Peter Dreisiger, HFRD Jan 1997
; Calculates the change in frequency required to cause a change in
; polarisation angle of dP, using the linearised model.
;
;=====
; input: freq = operating frequency (MHz)
;         fc = critical frequency (MHz)
;         dp = change in polarisation angle (rads)
;         dist = ground range (meters)
;         bearing = bearing (degrees east of north)
;=====
freq = double(freq)
fc = double(fc)
dP = double(dP)
dist = double(dist)
bearing = double(bearing)
magnetic, 132.25, -17.9, dip, decl, fH
fgm = double(0.5 * fH * cos(dip * !dtor) * Cos((decl -bearing) * !dtor) )

fo = freq +fgm
fx = freq -fgm
b0o = findb0(fo /fc, dist, findbmax(fo /fc) )
b0x = findb0(fx /fc, dist, findbmax(fx /fc) )
po = phasepath(fo /fc, b0o)
pdo = grouppath(fo /fc, b0o)
px = phasepath(fx /fc, b0x)
pdx = grouppath(fx /fc, b0x)
RETURN, 300d0 * dp / (atan(0d0, -1d0) *(pdo -po -pdx +px))
END
```

```
PRO TestError
; Test procedure used to evaluate the error in predicting a polarisation
; change of ANGLE using the linearised model
angle = !pi /2
bearing = 0.0
cnt = 0
err = fltarr(1001)
smoth = fltarr(1001)
FOR FREQ=20d0, 10d0, -0.01 DO BEGIN
  print bar, cnt, 1000
  cnt = cnt +1
  df = dfdp(freq, 9, angle, 2000000, bearing)
  err((freq -10) * 100) = polarrotation(freq +df, 2000000, bearing, 9) $
  - polarrotation(freq, 2000000, bearing, 9) -angle
ENDFOR
FOR I=2, 998 DO BEGIN
  smoth(i) = 0.2 * (err(i -2) +err(i -1) +err(i) +err(i +1) +err(i +2) )
ENDFOR

stop
END
```

```
PRO Distband
; Procedure to calculate the polarisation ground range
fc = double(8.97770275)
b = fltarr(361, 100)
read, freq, PROMPT='Enter operating frequency: '
freq = double(freq)
```

```
FOR DIST=1500000d0, 2500000d0, 10000 DO BEGIN
  print bar, (dist -1500000) /10000, 100
  fr = 0.8d0 * sqrt(1d0 +(dist /grouppath(0.8, findb0(0.8, dist, $
  findbmax(0.8) ) ) ) ^2)
  FOR BEAR=-45d0, 45d0, 0.05 DO BEGIN
    b((bear +45) * 4, (dist -1500000) /10000) = abs(cos(polarrotation(fr * fc,
    $
    dist, bear, fc) ) )
  ENDFOR
ENDFOR
stop
END
```

```
PRO Polarband
; Plots the polarisation bandwidth as a function of frequency
; for a given bearing.
fc = double(8.97770275)
b = dblarr(21, 5)
bearing = 0.0
FOR FREQ=double(10), 20, 0.5 DO BEGIN
  print, freq
  b((freq -10) * 2, 0) = dfdp(freq, fc, !pi /2, 1750000, 0)
  b((freq -10) * 2, 1) = dfdp(freq, fc, !pi /2, 1750000, 15)
  b((freq -10) * 2, 2) = dfdp(freq, fc, !pi /2, 1750000, 30)
  b((freq -10) * 2, 3) = dfdp(freq, fc, !pi /2, 1750000, 45)
  b((freq -10) * 2, 4) = dfdp(freq, fc, !pi /2, 1750000, 60)
ENDFOR
stop
END
```

```
FUNCTION Findb0, fR, dist, beta0max
; author: Peter Dreisiger, IE HFRD DSTO 96-97
; -----
; This function returns the value of b0 for which a ray
; of frequency fR=f/c will cover a ground distance of
; dist. beta0max is a parameter found using 'findbmax'
; that specifies at what elevation angle the high-angle
; ray commences.
; If the specified ground range is not possible at the
; given frequency, -1 is returned.
fR = double(fR)
dist = double(dist)
beta0max = double(beta0max)
accuracy = 0.002
bbot = double(0.0)
btop = beta0max
dmx = groundrange(fR, bbot)
dnn = groundrange(fR, btop)
diff = dmx -dnn

IF ((dist GE dnn) AND (dist LE dmx) AND (beta0max GE 0) ) THEN BEGIN
  WHILE (diff GT accuracy) DO BEGIN
    dmd = groundrange(fR, 0.5 *(bbot +btop) )
    diff = abs(dist -dmd)
    IF ((dist GE dnn) AND (dist LE dmd) ) THEN bbot = 0.5 *(btop +bbot)
    IF ((dist GE dmd) AND (dist LE dmx) ) THEN btop = 0.5 *(btop +bbot)
    dmx = groundrange(fR, bbot)
    dnn = groundrange(fR, btop)
  ENDWHILE
  b0 = 0.5 *(bbot +btop)
ENDIF ELSE BEGIN
  b0 = -1
ENDELSE
RETURN, b0
END
```

```
FUNCTION Findbmax, fR
; author: Peter Dreisiger, IE HFRD DSTO 96-97
; -----
; This function returns the maximum value of b0 for which the
; ray is a low-angle wave
fR = double(fR)
last = double(groundrange(fR, 0) )
IF (last GT 0.0) THEN BEGIN
  current = double(last)
  beta = 0.0
  betamax = 90 * !dior
  WHILE ((beta LE 90) AND (BETAMAX EQ 90 * !dior) ) DO BEGIN
    current = groundrange(fR, beta * !dior)
    IF (current GT last) THEN betamax = double(beta * !dior)
    last = current
    beta = beta +0.5
  ENDWHILE
ENDIF ELSE BEGIN
  betamax = -1
ENDELSE
RETURN, betamax
END
```

```
FUNCTION GroundRange, fR, b0
; author: Peter Dreisiger, IE HFRD DSTO 96-97
; -----
; Uses Croft's solution for ray propagation in a QP concentric
; ionosphere to determine the ground range for a given frequency
```

```
; ratio (f/c) and elevation angle.
ro = double(6370000)
rm = double(ro +300000)
ym = double(100000)
rb = double(rm -ym)
fR = double(fR)
b0 = double(b0)
d = 0.0

a = 1.0 -1.0 /fR^2 +(rb /(fR * ym) ) ^2
b = -2.0 * rm *(rb /(fR * ym) ) ^2
c = (rb * rm /(fR * ym) ) ^2 -(ro * cos(b0) ) ^2
IF ((b^2 -4 * a * c) GE 0) AND (b0 GE 0) ) THEN BEGIN
  g = acos(ro /rb * cos(b0) )
  d = 2 * ro *((g -b0) -ro * cos(b0) * 0.5 /sqrt(c) * alog((b^2 -4 * a * c) $
    / (4.0 * c *(sin(g) +sqrt(c) /rb +0.5 * b /sqrt(c) ) ^2) ) )

ENDIF
IF (d LT 0) THEN d = 0
RETURN, double(d)
END
```

```
FUNCTION Grouppath, fR, b0
; author: Peter Dreisiger, IE HFRD DSTO 96-97
; -----
; Uses Croft's solution for ray propagation in a QP concentric
; ionosphere to determine the group path for a given frequency
; ratio (f/c) and elevation angle.
ro = double(6370000)
rm = double(ro +300000)
ym = double(100000)
rb = double(rm -ym)
fR = double(fR)
b0 = double(b0)
p = 0.0

a = 1.0 -1.0 /fR^2 +(rb /(fR * ym) ) ^2
b = -2.0 * rm *(rb /(fR * ym) ) ^2
c = (rb * rm /(fR * ym) ) ^2 -(ro * cos(b0) ) ^2
g = acos(ro /rb * cos(b0) )
IF ((b^2 -4 * a * c) GE 0) THEN $
  p = 2 *(rb * sin(g) -ro * sin(b0) +1 /a *( -rb * sin(g) -0.25 * b /sqrt(a) $
    * alog((b^2 -4 * a * c) / (2 * a $
      * rb +b +2 * rb * sqrt(a) * sin(g) ) ^2) ) )

RETURN, double(p)
END
```

```
FUNCTION Phasepath, fR, b0
; author: Peter Dreisiger, IE HFRD DSTO 96-97
; -----
; Uses Croft's solution for ray propagation in a QP concentric
; ionosphere to determine the phase path for a given frequency
; ratio (f/c) and elevation angle.
ro = double(6370000)
rm = double(ro +300000)
ym = double(100000)
rb = double(rm -ym)
fR = double(fR)
b0 = double(b0)
p = 0.0
```

```
a = 1.0 -1.0 /fR^2 +(rb /(fR * ym) ) ^2
b = -2.0 * rm *(rb /(fR * ym) ) ^2
c = (rb * rm /(fR * ym) ) ^2 -(ro * cos(b0) ) ^2
```

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```
g = acos(ro /rb * cos(b0) )
IF ((b^2 -4 * a * c) GE 0) THEN $
P=2.0 * ( -ro * sin(b0) +b /4.0 *(1.0 /sqrt(a) * alog((b^2 -4.0 * a * c) $
/ (4.0 *(a * rb +b /2.0 +sqrt(a) * rb * sin(g) )^2) ) $
+rm /sqrt(c) * alog((b^2 -4.0 * a * c) / (4.0 * c *( $
sin(g) +sqrt(c) /rb +0.5 * b /sqrt(c) )^2 ) ) ) )
RETURN, double(p)
END
```

```
PRO Magnetic, Glong, Glat, Dip, Decl, fh
; author: Chris Coleman, IE HFRD DSTO 94
; -----
; Calculates the magnetic field wuantites for an offset dipole
; Glat = mean geographic latitude
; glong = mean geographic longitude
;
; returns Dip, Decl, fh
;
; Fgm=0.5*fH*cos(dip*!dtor)*cos(decl-bearing)*!dtor)
scale = 45.0 /atan(1.0)
Dlong = double(295.75)
Dlat = double(77.75)
SGlat = sin(Glat /scale)
SDlat = sin(Dlat /scale)
CGlat = cos(Glat /scale)
CDlat = cos(Dlat /scale)
Clong = cos((Glong -Dlong) /scale)
Slong = sin((Glong -Dlong) /scale)
Smlat = SGlat * SDlat +CGlat * CDlat * Clong
Cmlat = sqrt(abs(1.0 -Smlat^2) )
Dip = scale * atan(Smlat, Cmlat)
glo = -CDlat * Slong
gla = CGlat * SDlat -SGlat * CDlat * Clong
Decl = scale * atan(glo, gla)
fh = 0.73 * sqrt(1.0 +3.0 * Smlat^2)
END
```